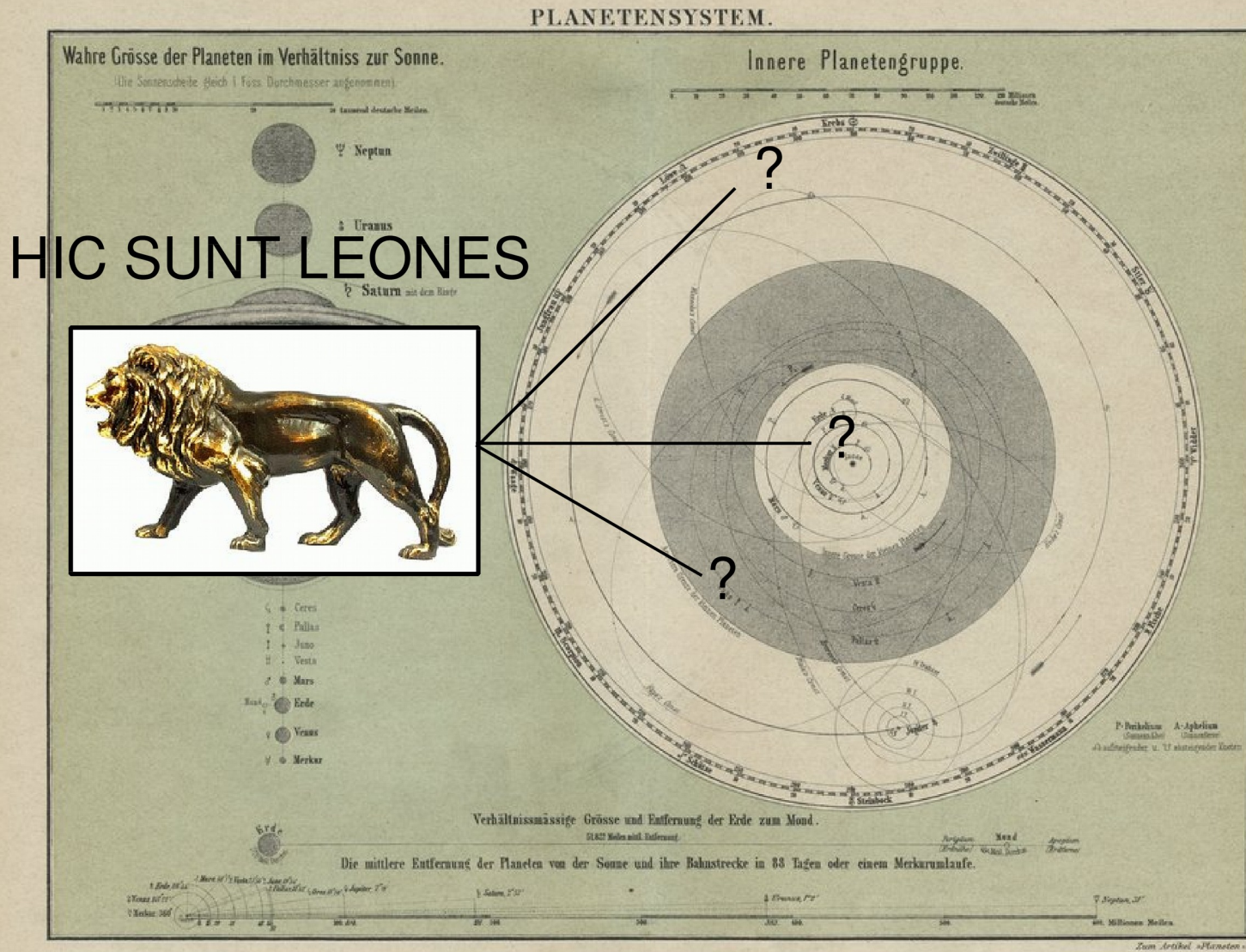
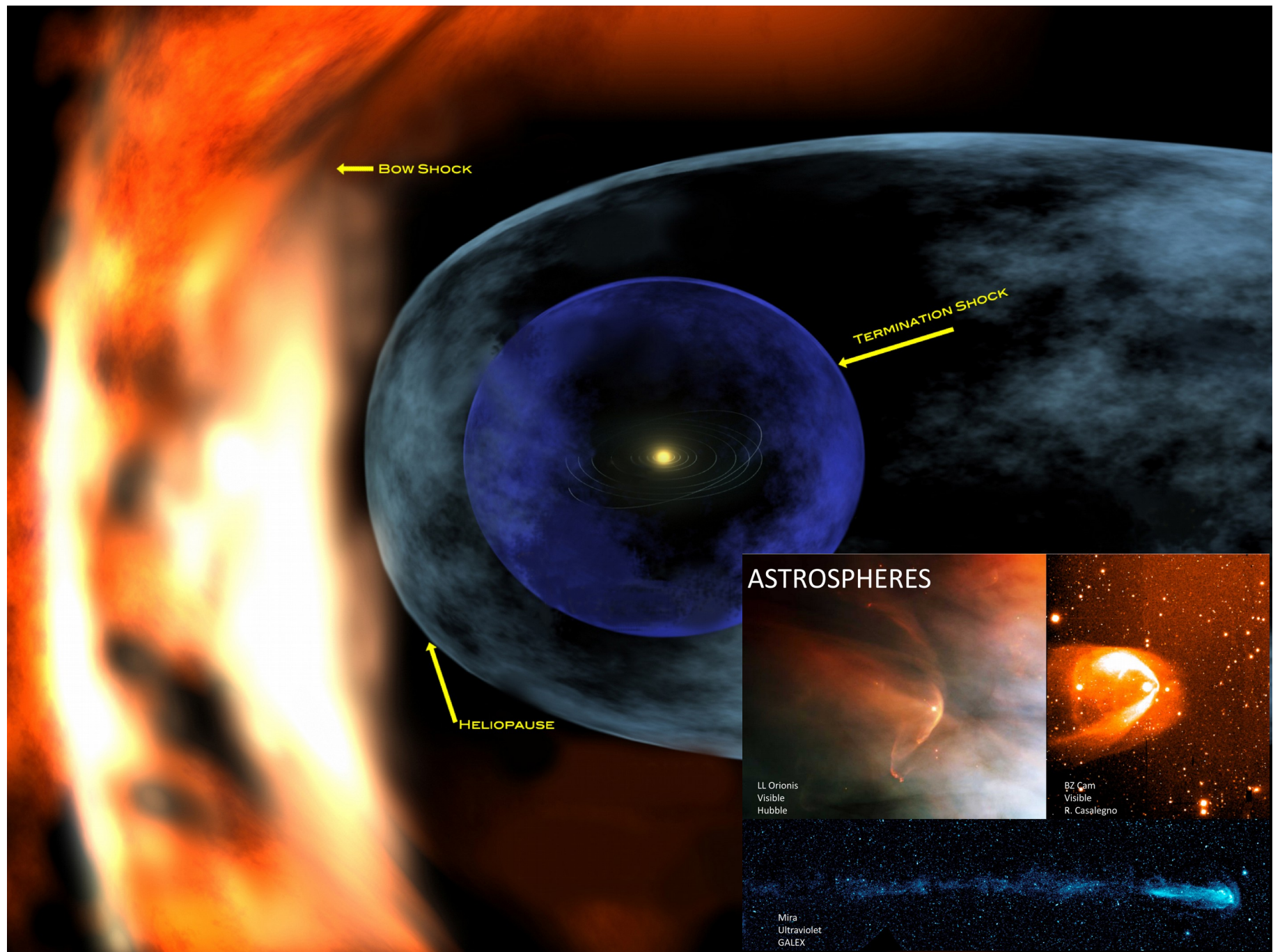


The Solar Wind

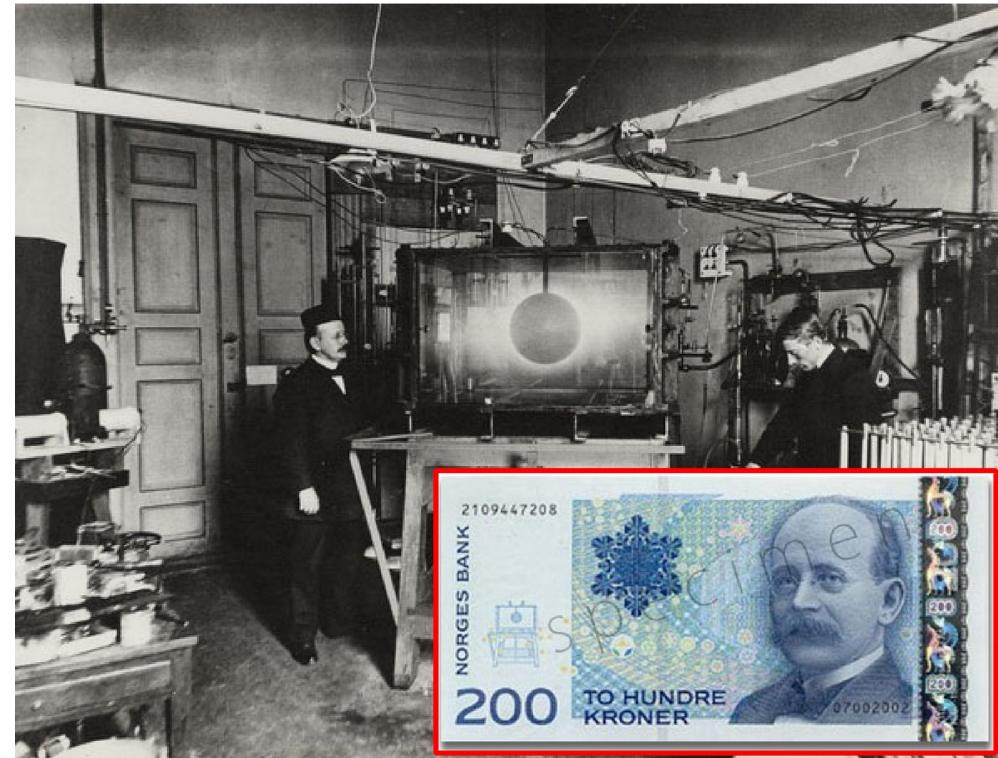
L. Berger

Extraterrestrial Physics, Institute for Experimental and Applied Physics, Christian-Albrechts-University Kiel, Germany

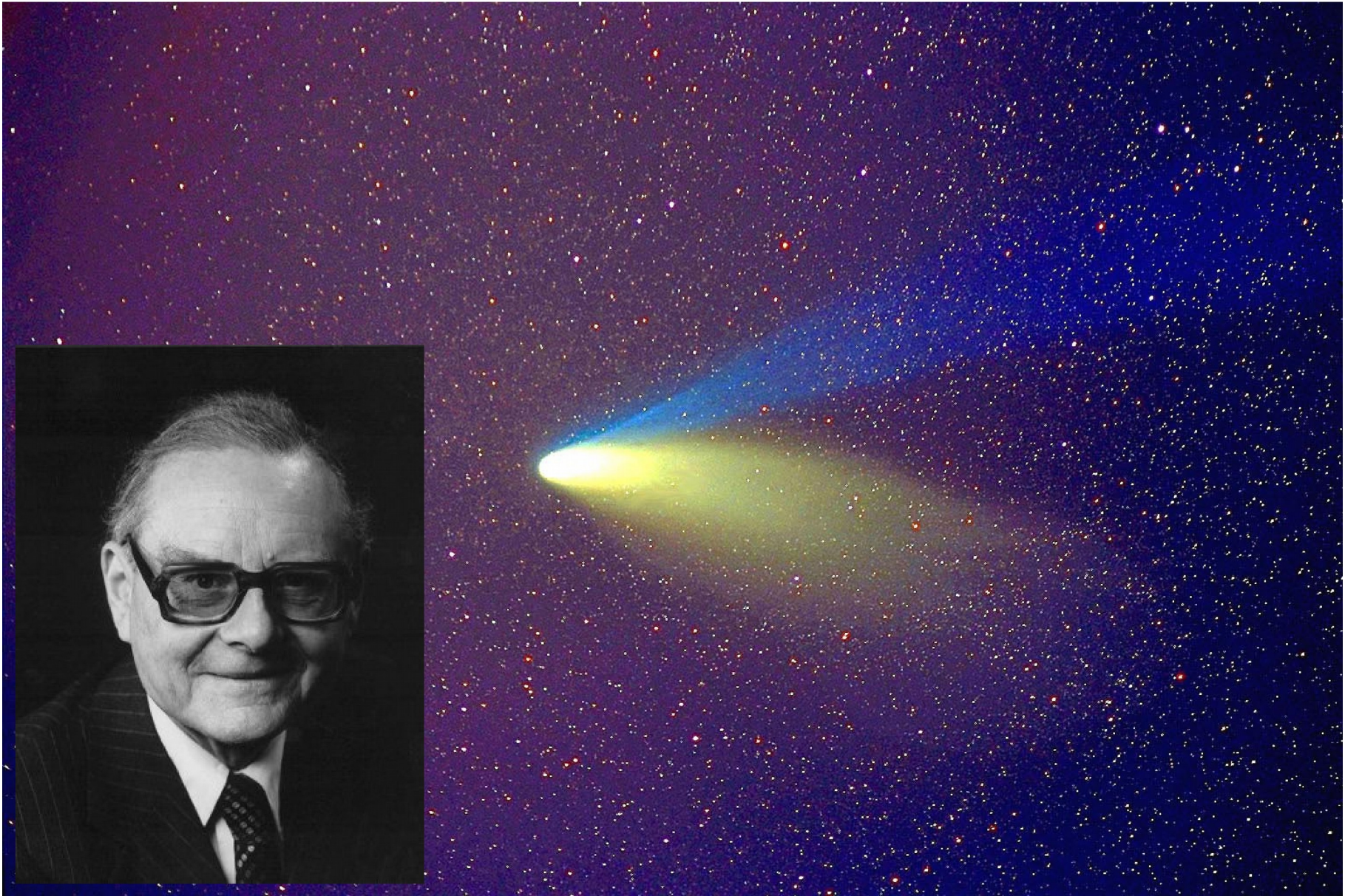




1859 : Richard Carrington – Related observed flare to geomagnetic storm
~1895 : Kristian Birkeland – Transient Solar ions and electrons



1951 : Ludwig Biermann – Suggests continuous Solar Wind



04.10.1957 : Soviet Sputnik 1 – Beginning of Space Age 1958 : Eugene Parker – Suggests Solar Wind model Predicts Parker Spiral

DYNAMICS OF THE INTERPLANETARY GAS AND MAGNETIC FIELDS*

E. N. PARKER

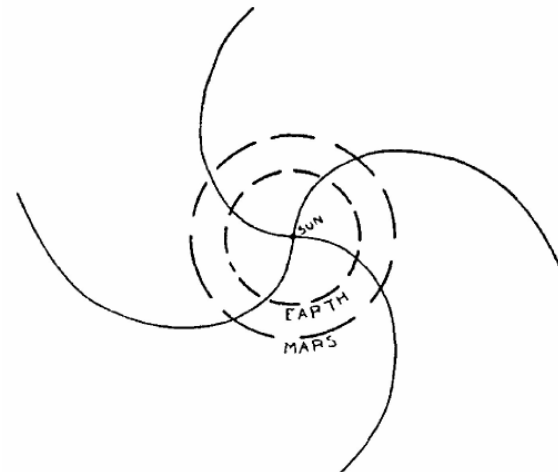
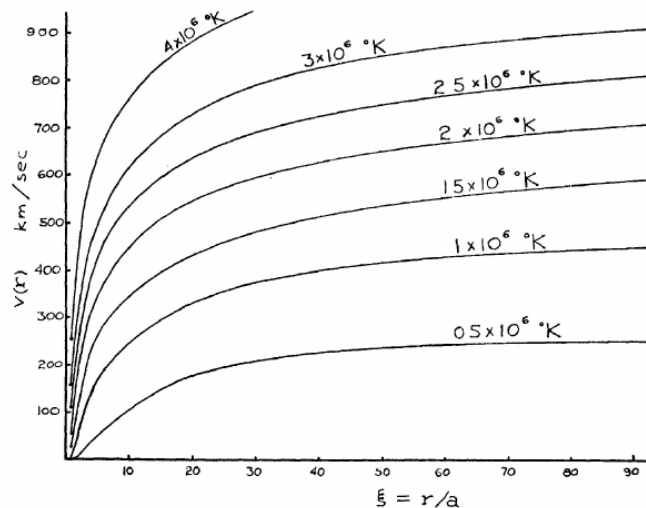
Enrico Fermi Institute for Nuclear Studies, University of Chicago

Received January 2, 1958

ABSTRACT



We consider the dynamical consequences of Biermann's suggestion that gas is often streaming outward in all directions from the sun with velocities of the order of 500–1500 km/sec. These velocities of 500 km/sec and more and the interplanetary densities of 500 ions/cm³ (10¹⁴ gm/sec mass loss from the sun) follow from the hydrodynamic equations for a 3×10^6 °K solar corona. It is suggested that the outward-streaming gas draws out the lines of force of the solar magnetic fields so that near the sun the field is very nearly in a radial direction. Plasma instabilities are expected to result in the thick shell of disordered field (10⁻⁵ gauss) inclosing the inner solar system, whose presence has already been inferred from cosmic-ray observations.



1959 : Soviet Lunar2 rocket confirmed stream of ions

THE STUDY OF INTERPLANETARY IONIZED GAS, HIGH-ENERGY ELECTRONS AND CORPUSCULAR RADIATION OF THE SUN, EMPLOYING THREE-ELECTRODE CHARGED PARTICLE TRAPS ON THE SECOND SOVIET SPACE ROCKET*

K. I. GRINGAUZ, V. V. BEZRUKIKH, V. D. OZEROV and R. E. RYBCHINSKII

Translated by R. MATTHEWS from *skusstvennye Sputniki Zemli*, 6, 101, (1961).

.....

3. Starting from 9.30 hr Moscow time on 13 September 1959 up to the moment of the container of the second space rocket reaching the moon the container was recorded as passing through a positive ion flux (in all probability protons) with energies exceeding 15 eV; $N \sim 2 \times 10^8 \text{ cm}^{-2}/\text{s}$. This follows from the fact that during this time approximately identical positive collector currents were recorded in all four traps (see the last section of Fig. 4.)

The existence at a different time of a proton flux with energies exceeding 25 eV was observed employing similar instrumentation at different distances from the earth (in particular with $R \sim 125,000 \text{ km}$) during a number of radiotelemetering data transmission sessions during the flight of the automatic interplanetary station in October 1959.

The proton fluxes recorded refer evidently to solar corpuscular radiation observed for the first time in this manner in interplanetary space outside the earth's magnetic field.

1962 : American Mariner2 measures speed and temperature

Science 7 December 1962:
Vol. 138 no. 3545 pp. 1095-1097
Solar Plasma Experiment
Marcia Neugebauer and Conway W. Snyder
Jet Propulsion Laboratory, California Institute of Technology, Pasadena

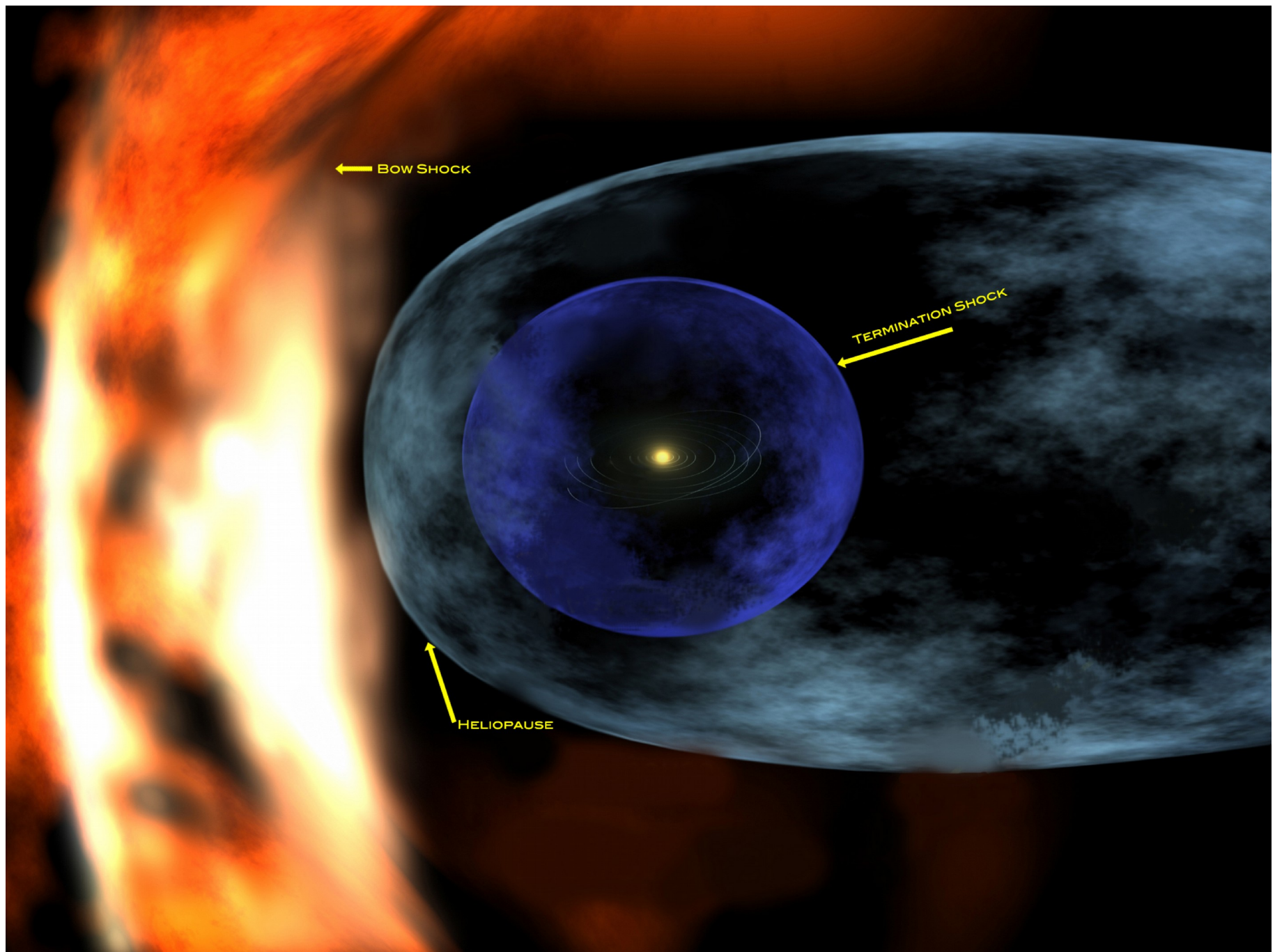
Solar Plasma Experiment

Abstract. A preliminary summary of the data received from the Mariner II solar plasma experiment for the period 29 August through 31 October 1962 is presented. During this period there was always a measurable flow of plasma from the direction of the sun. The velocity of this ion motion was generally in the range 400 to 700 km/sec. Time variations, plasma density, and ion temperatures are also discussed.

The Mariner II solar plasma experiment is made with a single electrostatic spectrometer which always points to within less than $\frac{1}{2}$ degree of the center of the sun. Positively

X. The measured velocity agrees fairly well with the value predicted from Parker's "solar wind" theory (4) but is higher than the value predicted from the observation of comet tail orientations (5) and much higher than the

3) The plasma energy density is much greater than the energy density of the magnetic field. Thus we may conclude that the magnetic field in interplanetary space is carried along by the plasma, the field giving little or no hindrance to the plasma flow.



Mean Plasma Parameter

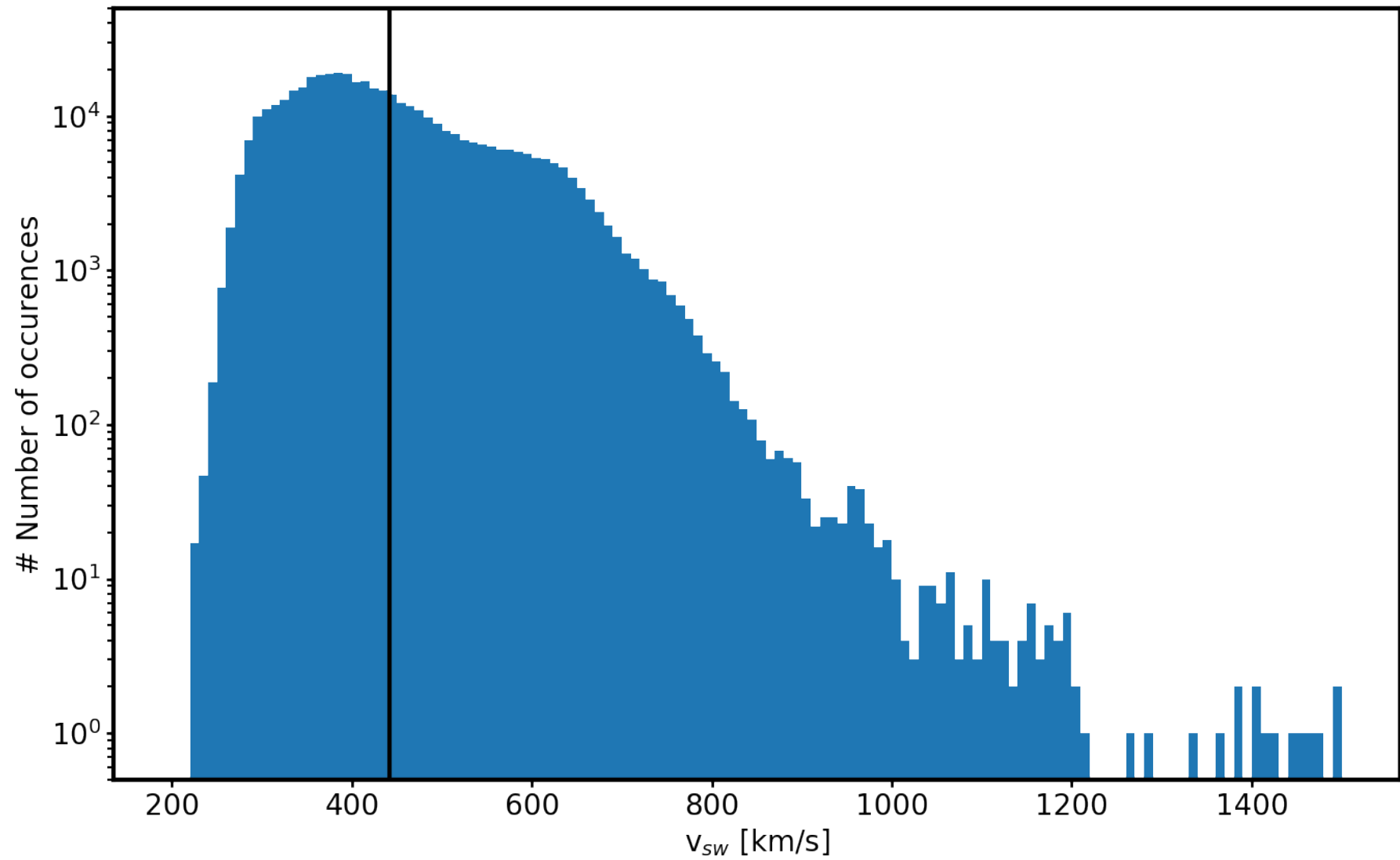
$$V_{\text{sw}} = 441 \pm 111 \text{ km/s}$$

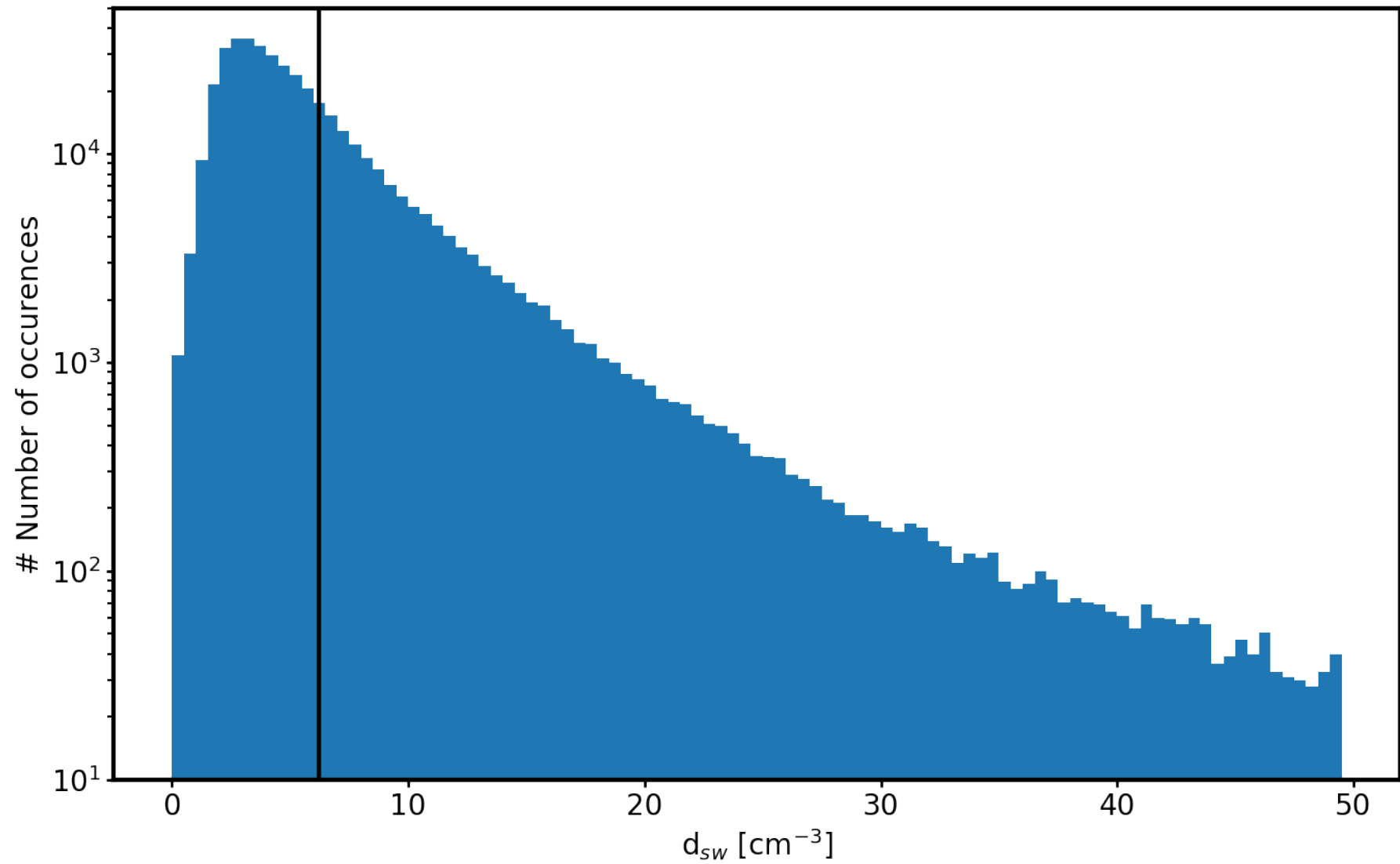
$$D_{\text{sw}} = 6.2 \pm 5.6 \text{ cm}^{-3}$$

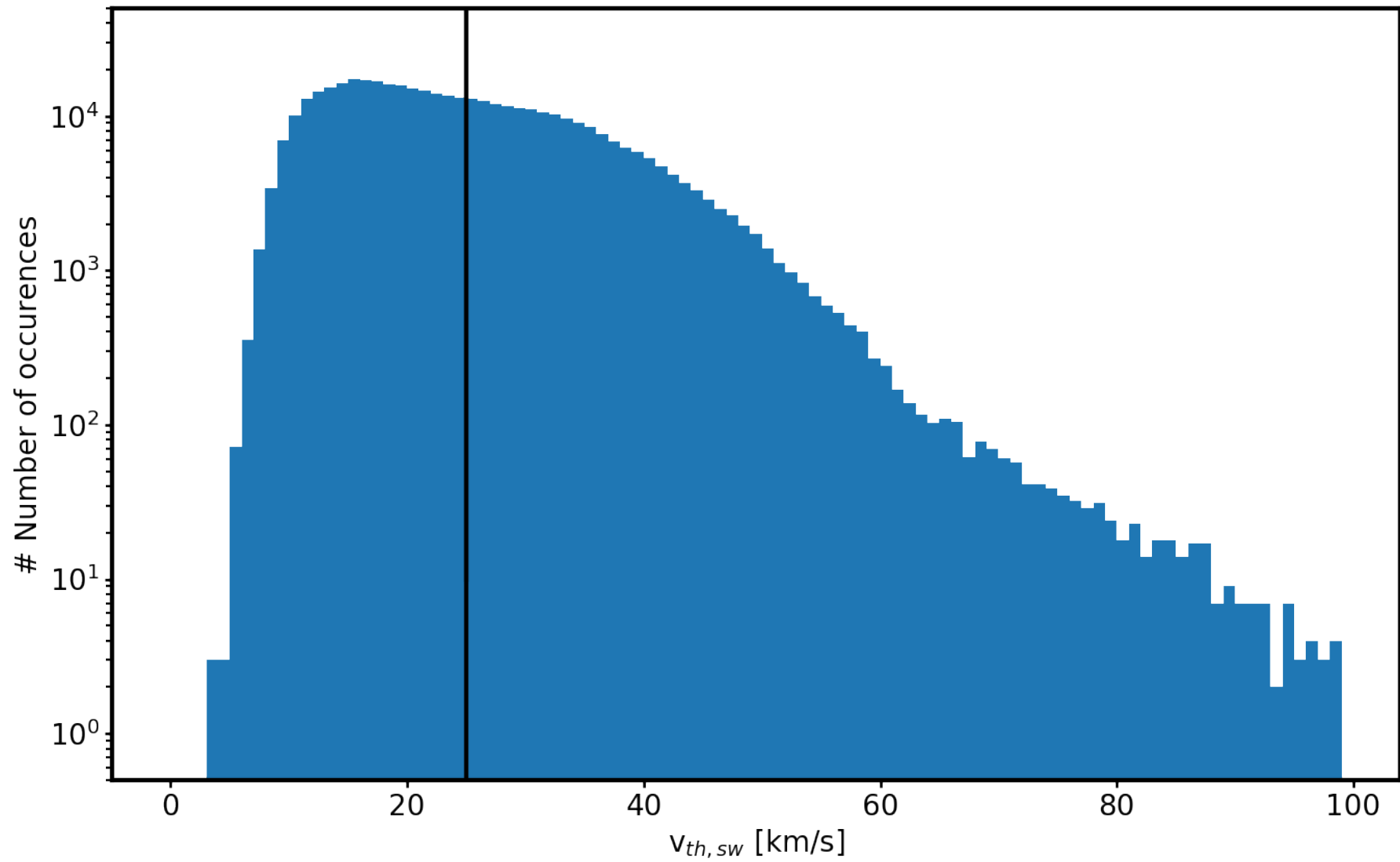
$$T_{\text{sw}} = 89 \cdot 10^3 \pm 84 \cdot 10^3 \text{ K}$$

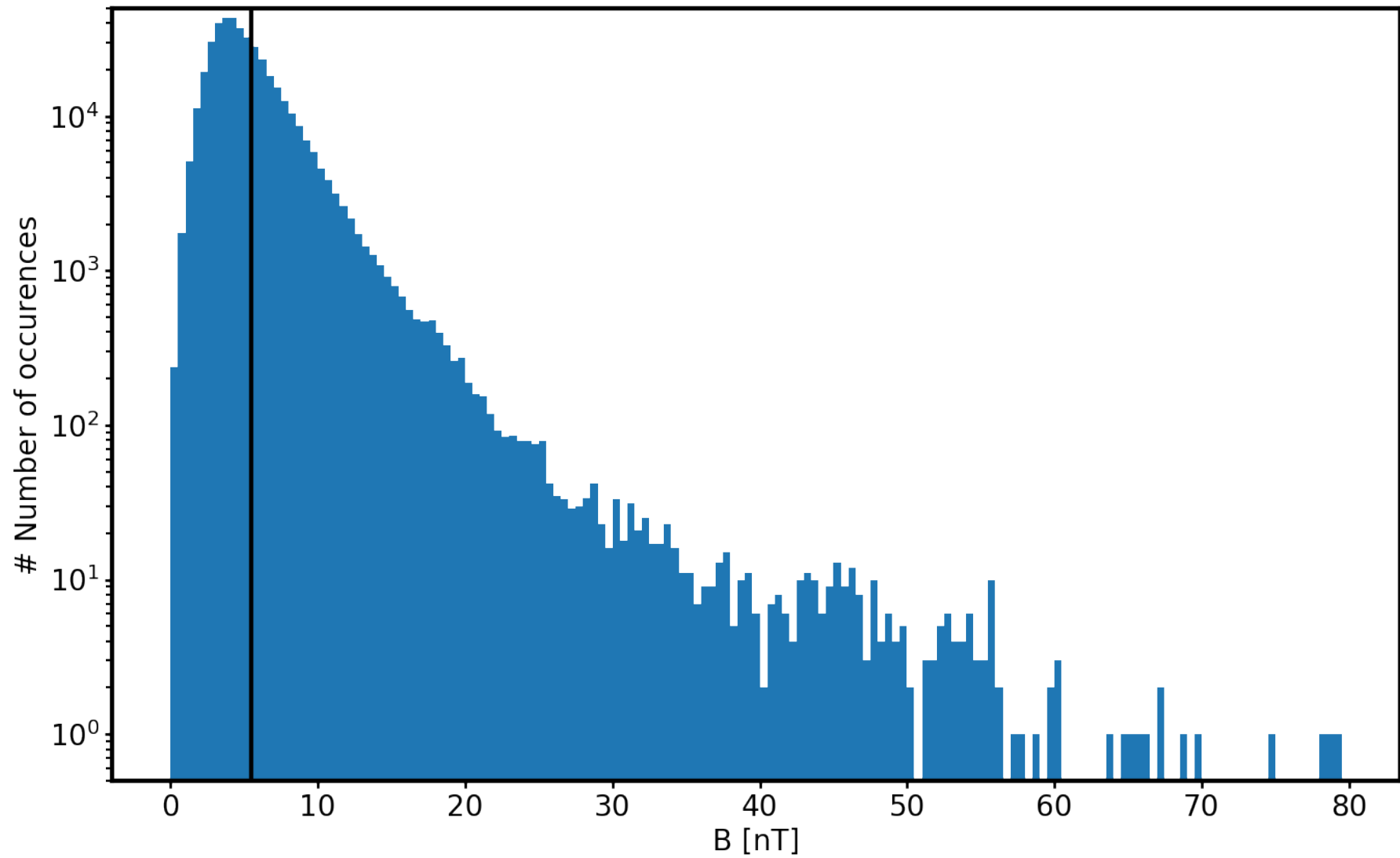
$$v_{\text{th,sw}} = 25 \text{ km/s} \pm 11 \text{ km/s}$$

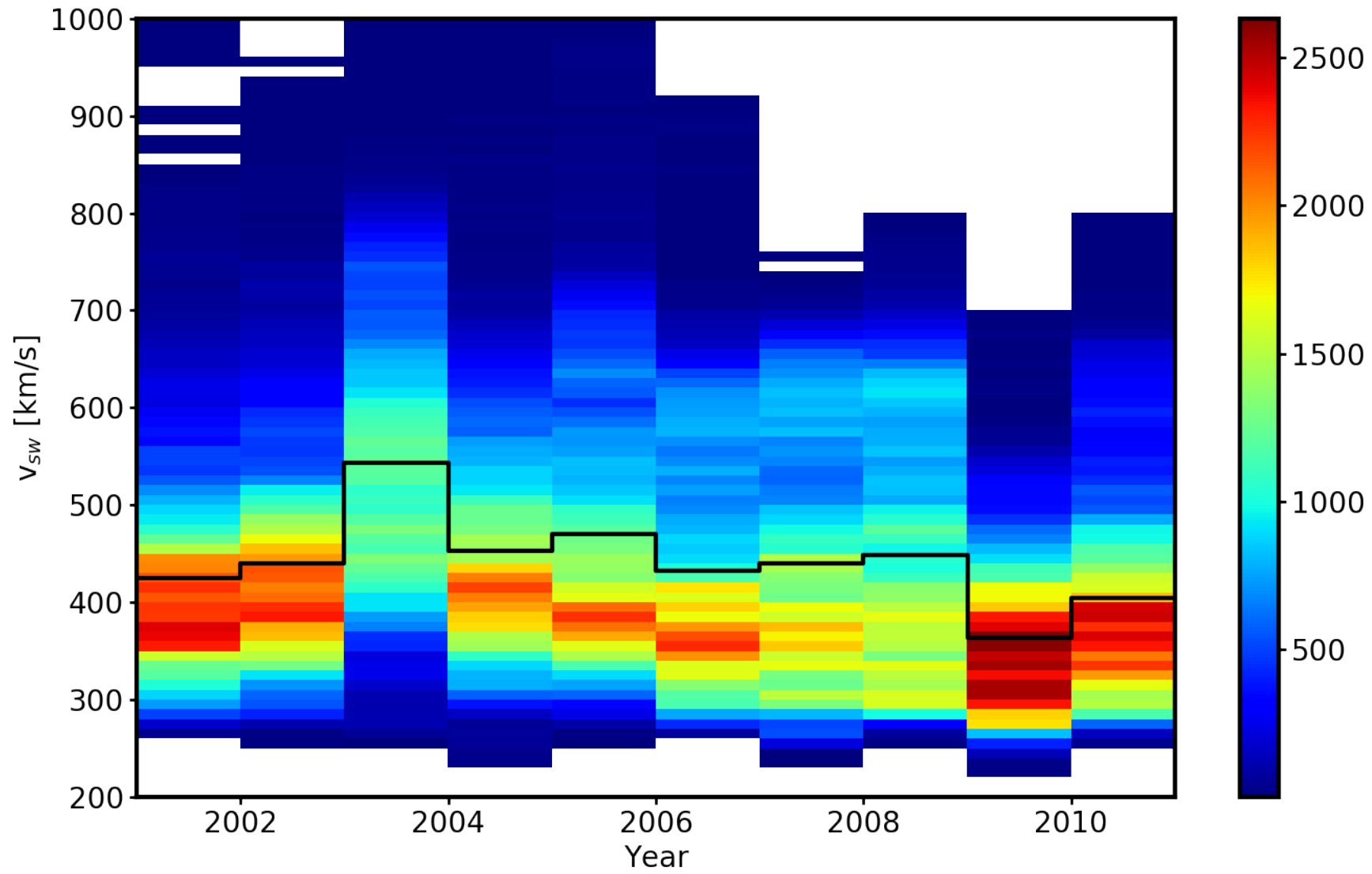
$$B = 5.45 \pm 3.2 \text{ nT}$$

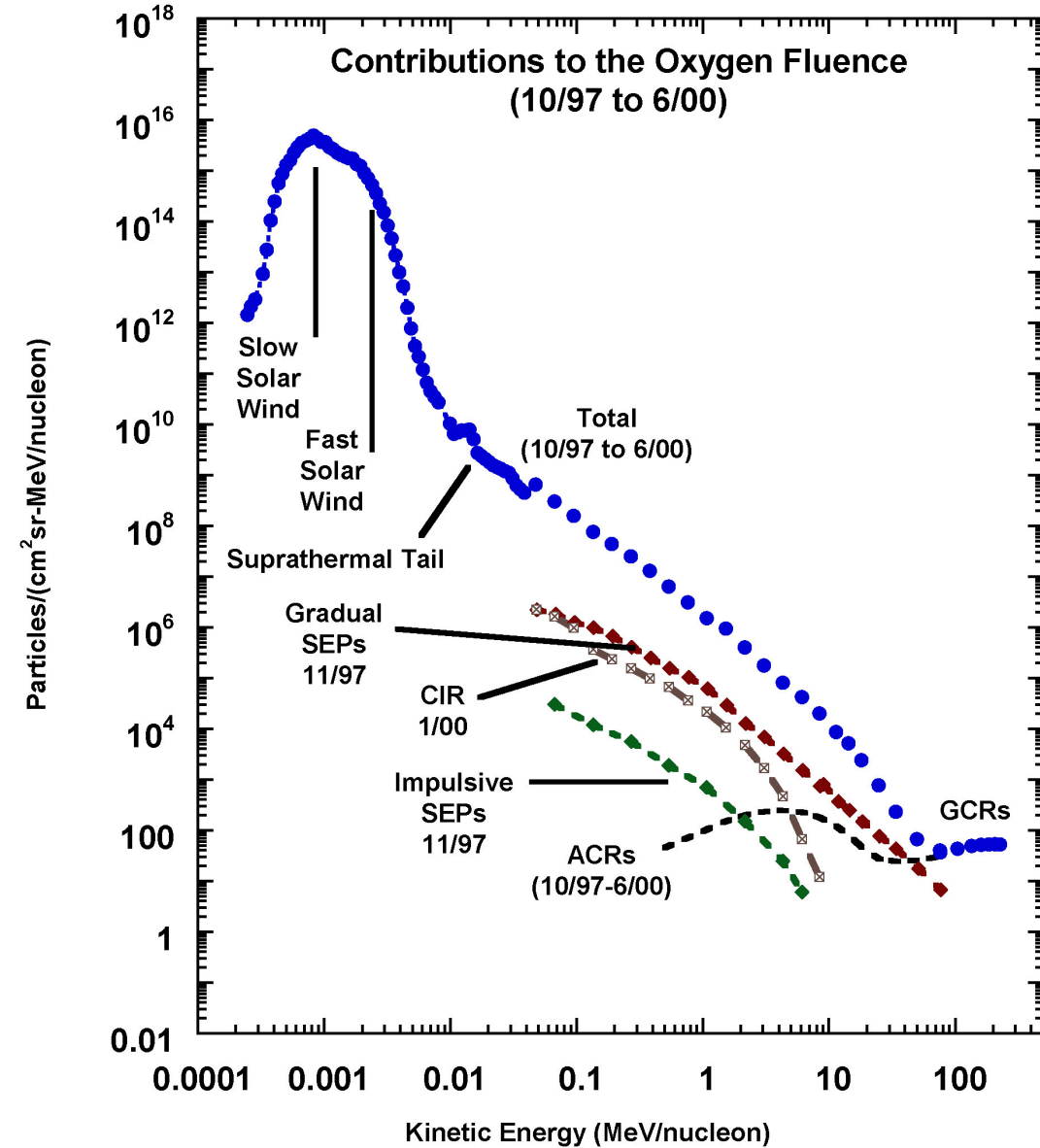
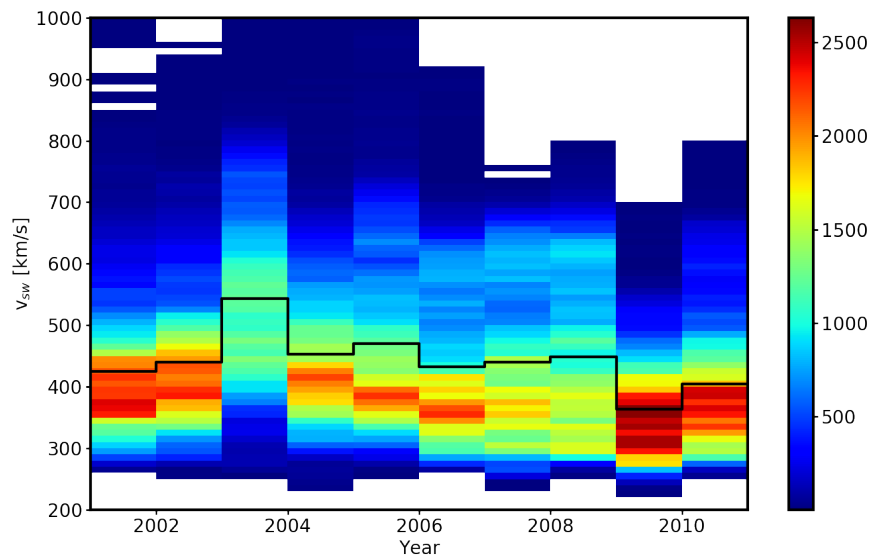
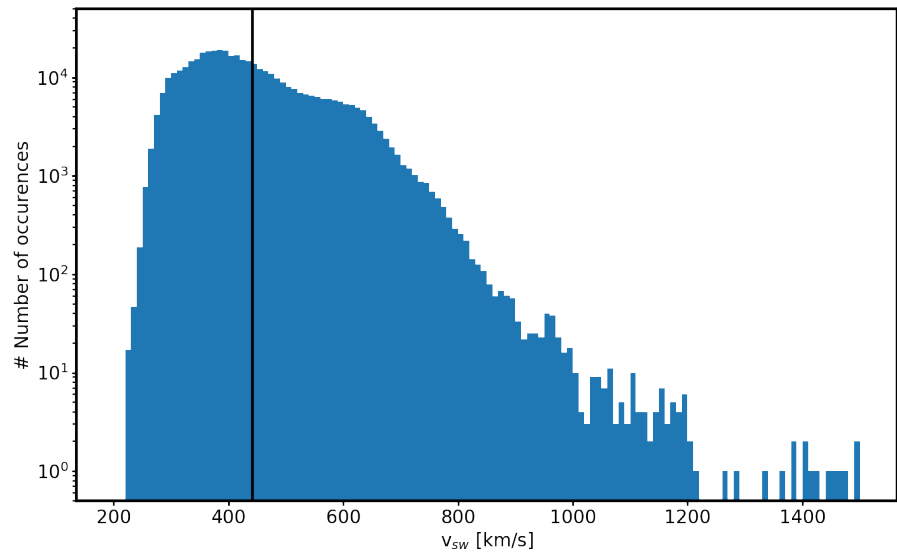




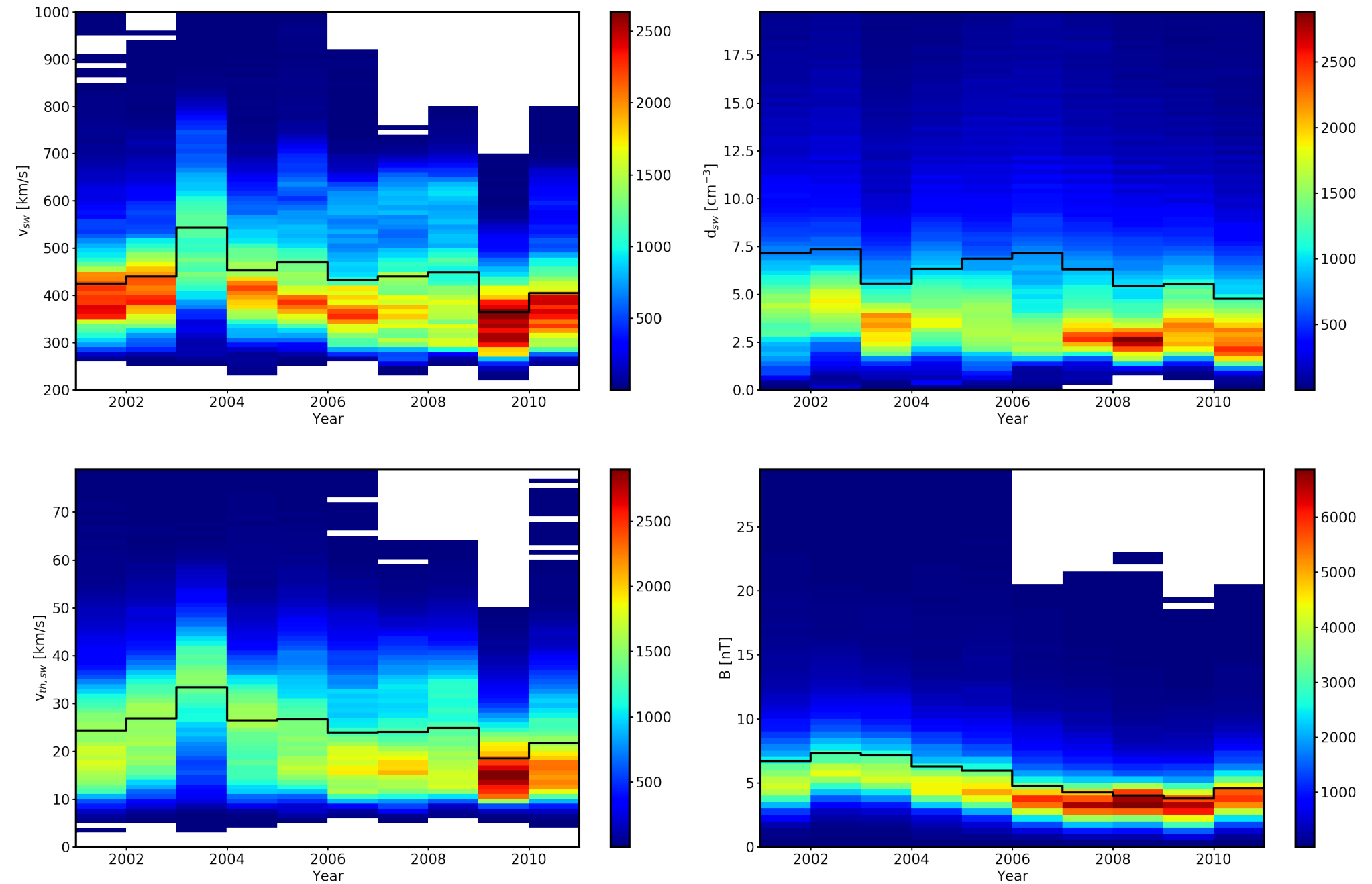


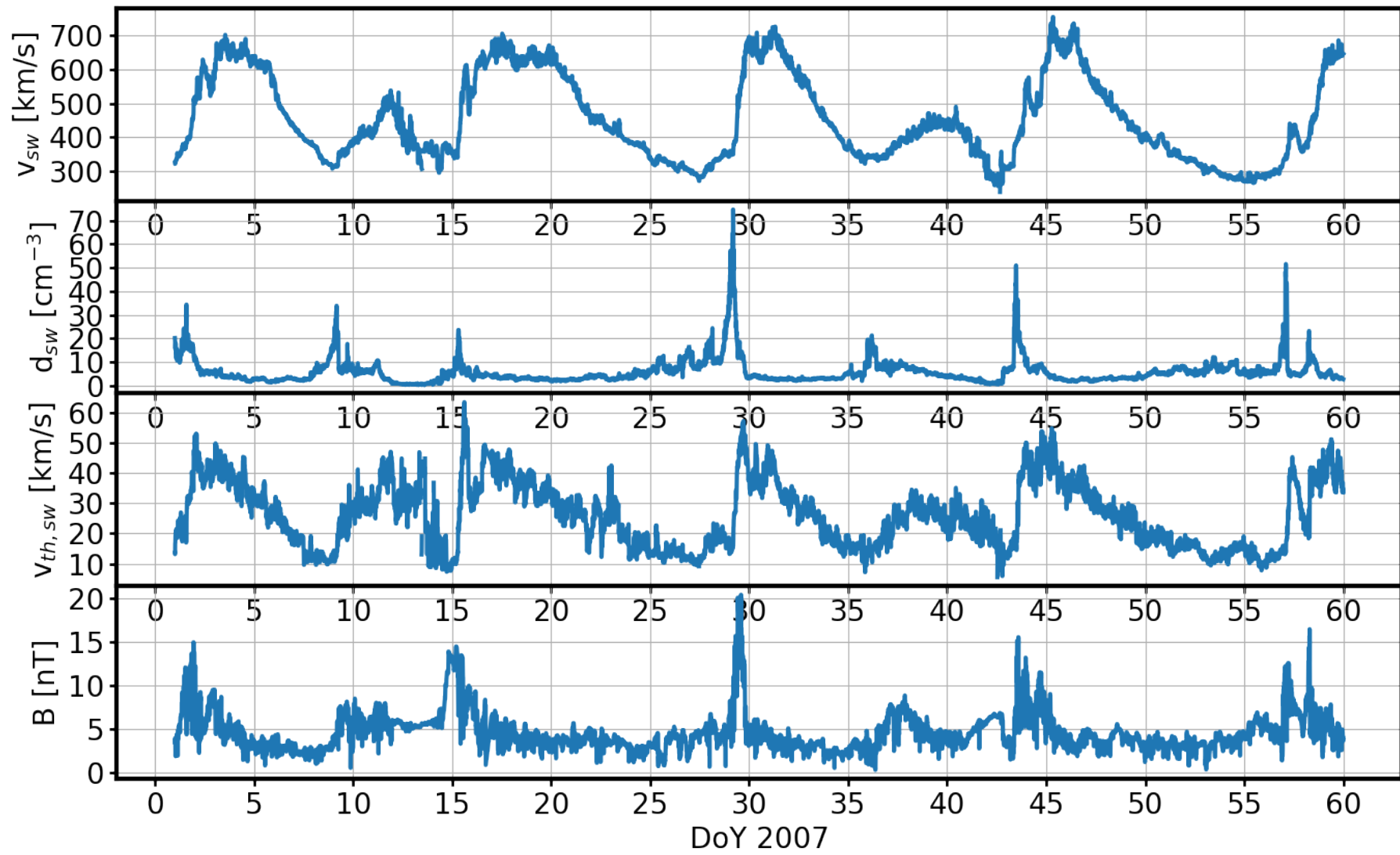


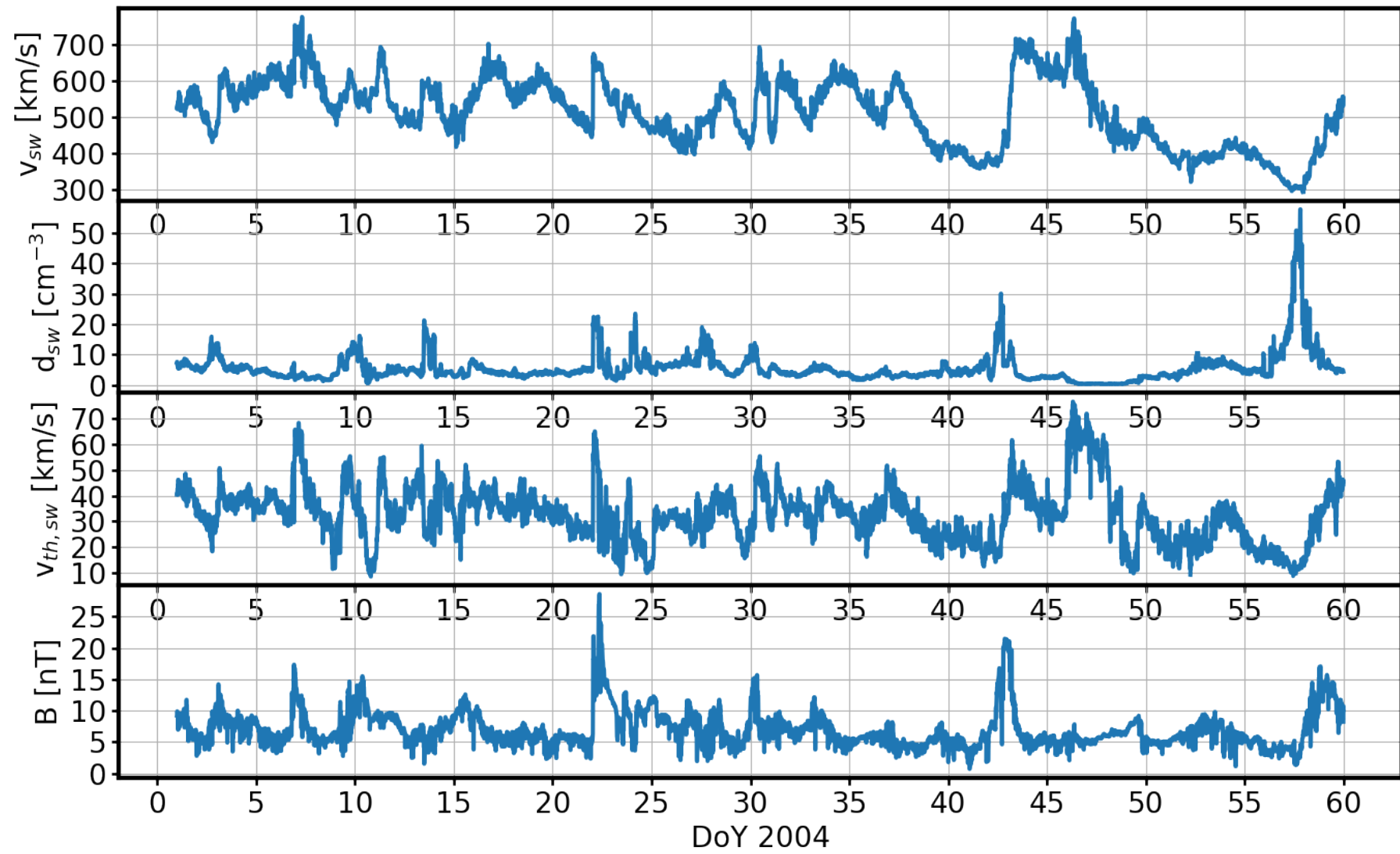


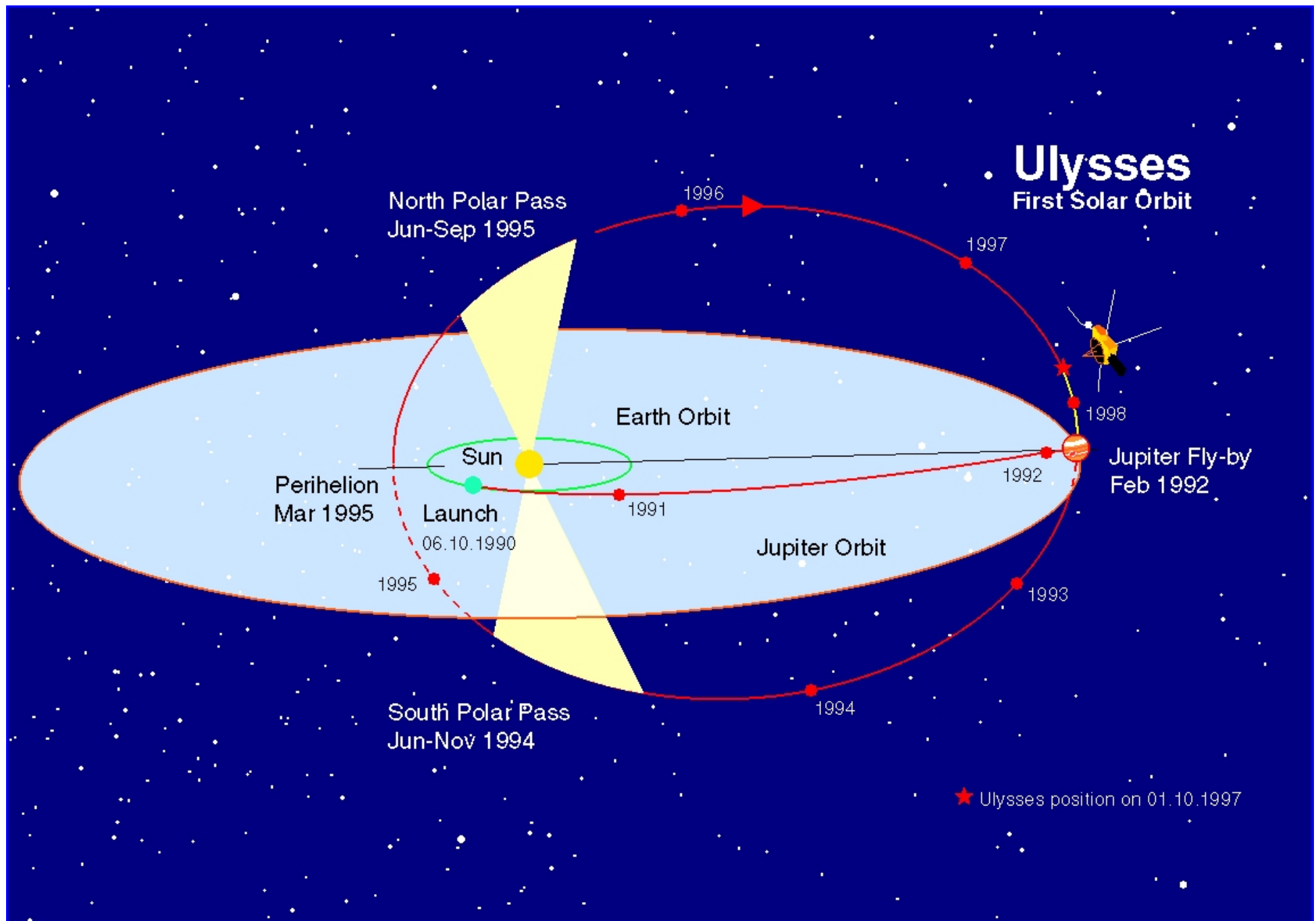


From <http://www.srl.caltech.edu>





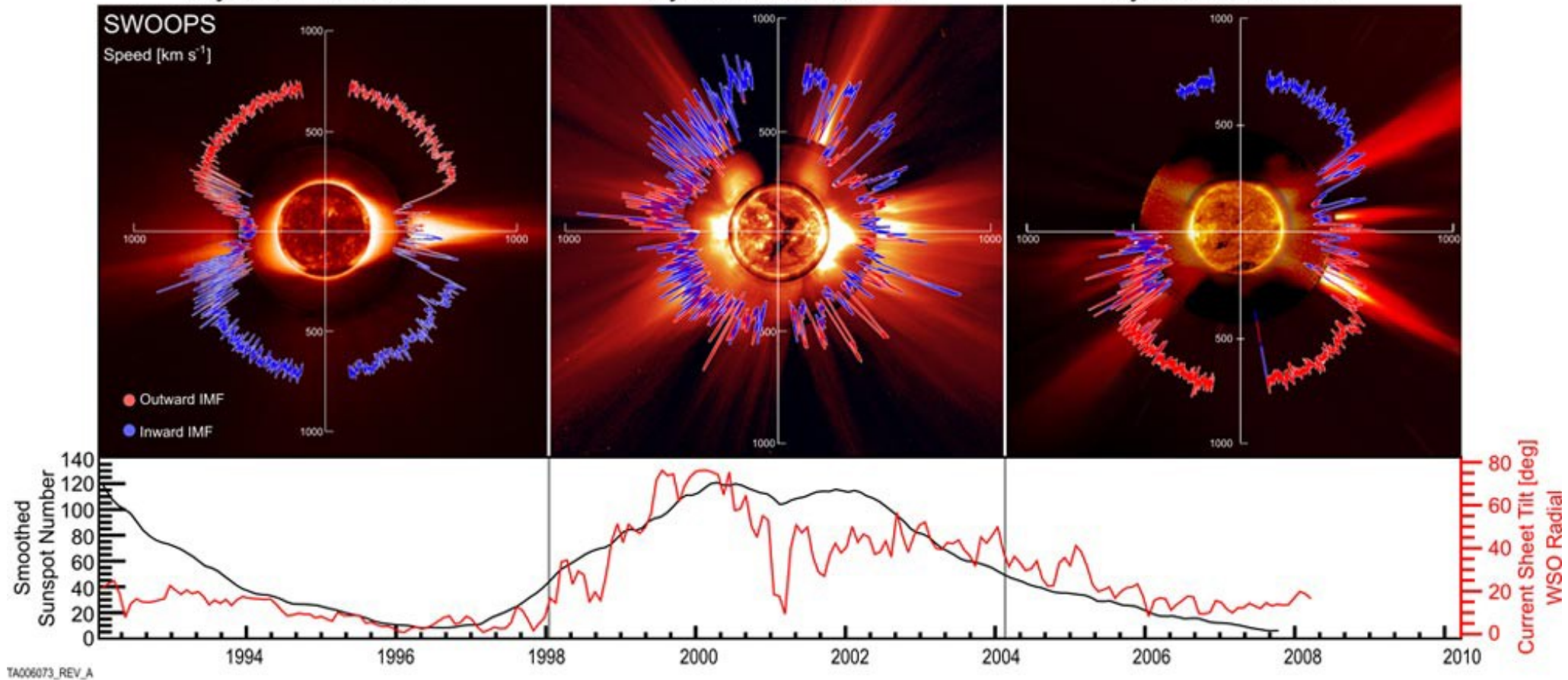


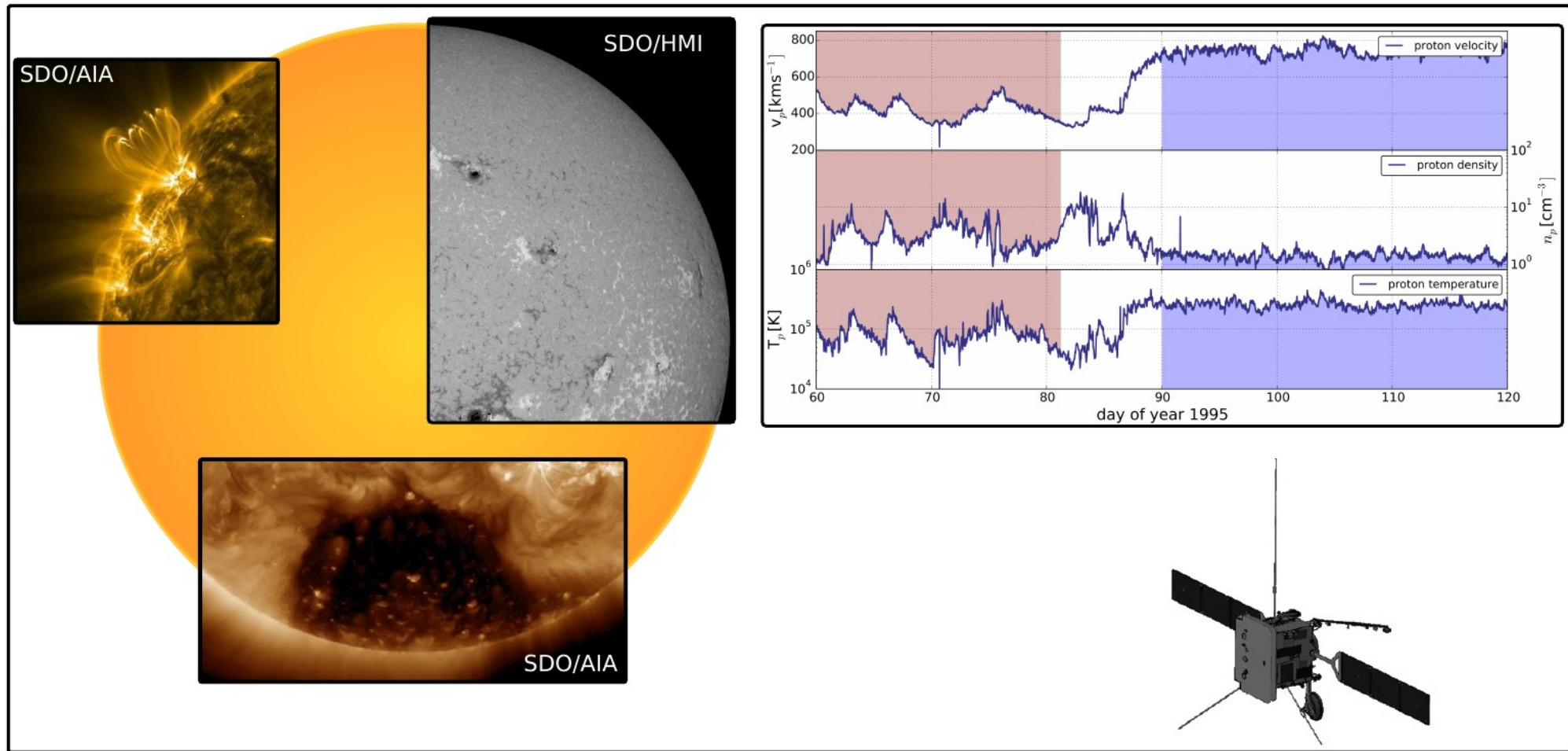


Ulysses First Orbit

Ulysses Second Orbit

Ulysses Third Orbit

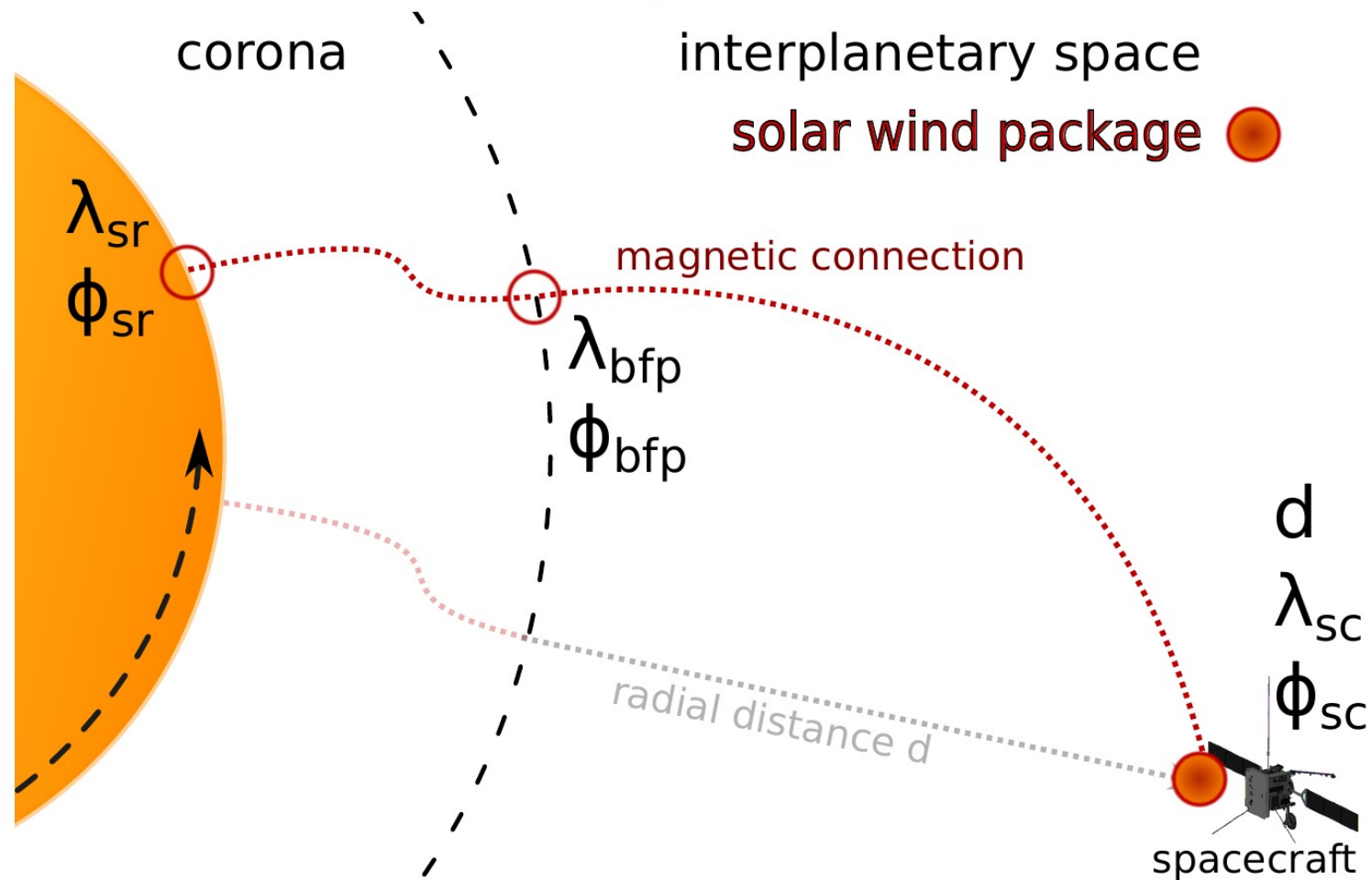




Source small pictures: Courtesy of NASA/SDO and the AIA and HMI science teams.

Linking in-situ Data to its Source

Back Mapping - Ballistic



$$\lambda_{sr}(d, v_{sw}) = \lambda_{sc} + \frac{\omega d}{v_{sw}} + \Delta \lambda_{mag}$$

$$\phi_{sr} = \phi_{sc} + \Delta \phi_{mag}$$

Linking in-situ Data to its Source

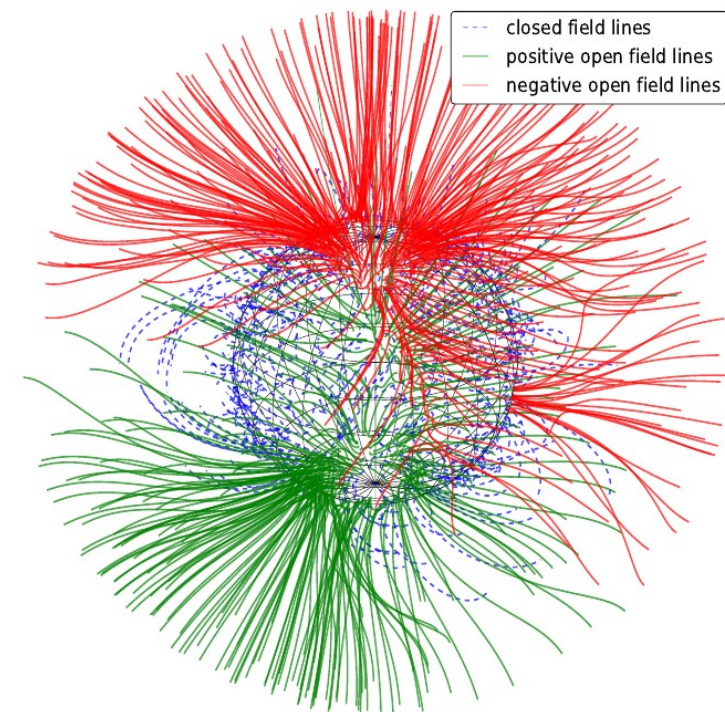
Back Mapping - Magnetic

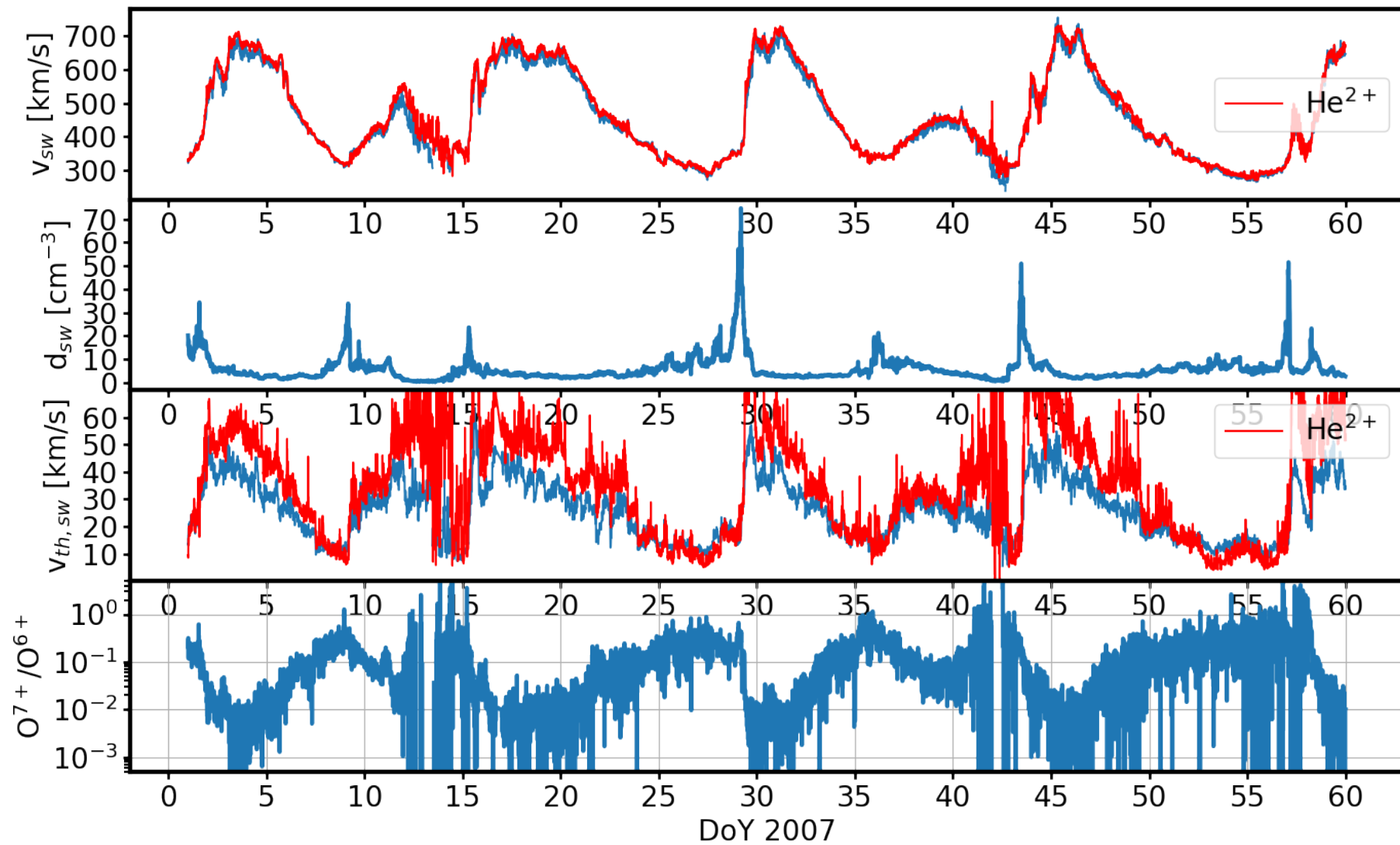
$\Delta\lambda_{mag}$ and $\Delta\phi_{mag}$ must be determined by using a model of the coronal magnetic field: The Potential Field Source Surface (PFSS) model (Altschuler & Newkirk (1969), here Kruse (2018)).

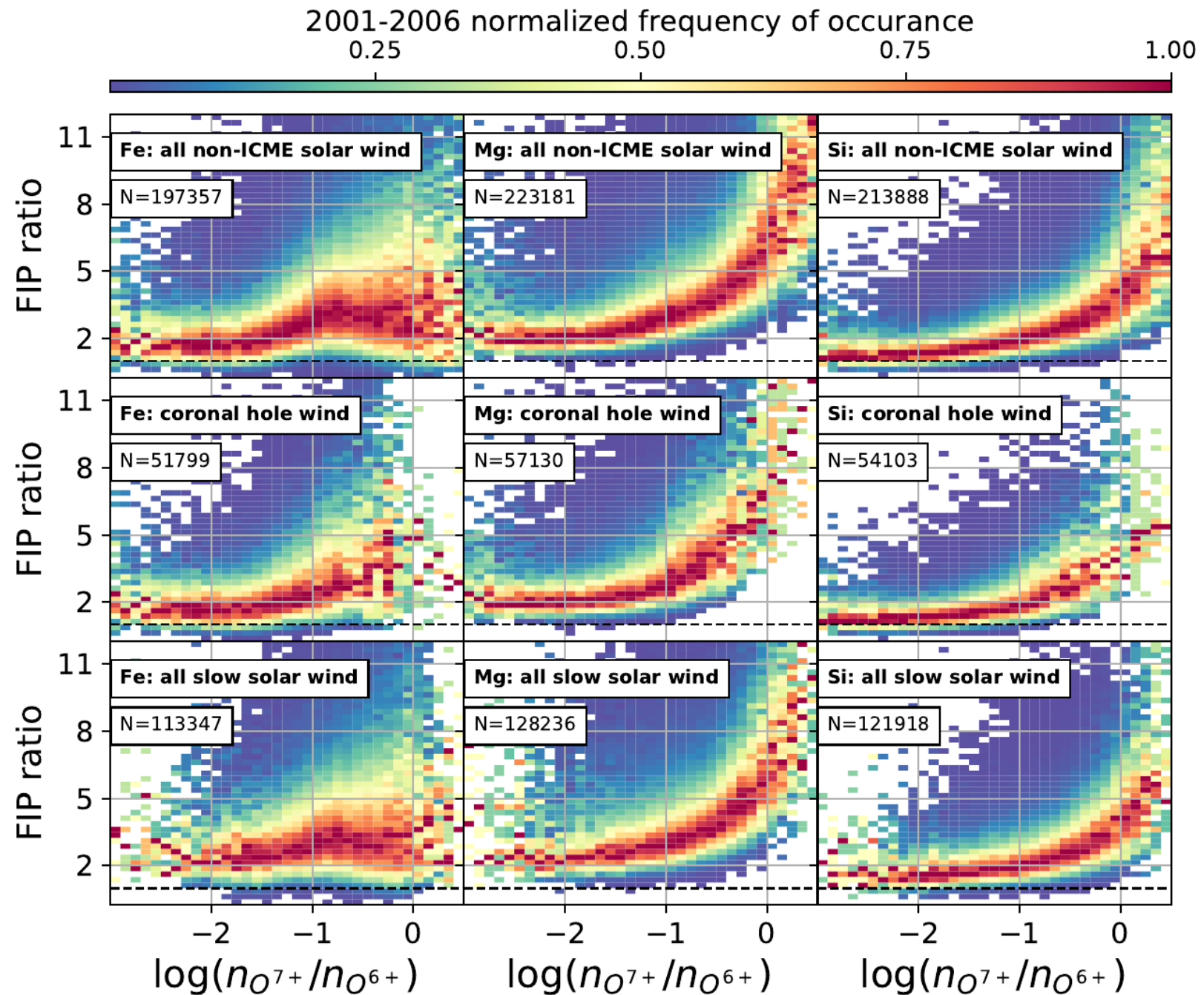
Assumption: a current-free corona

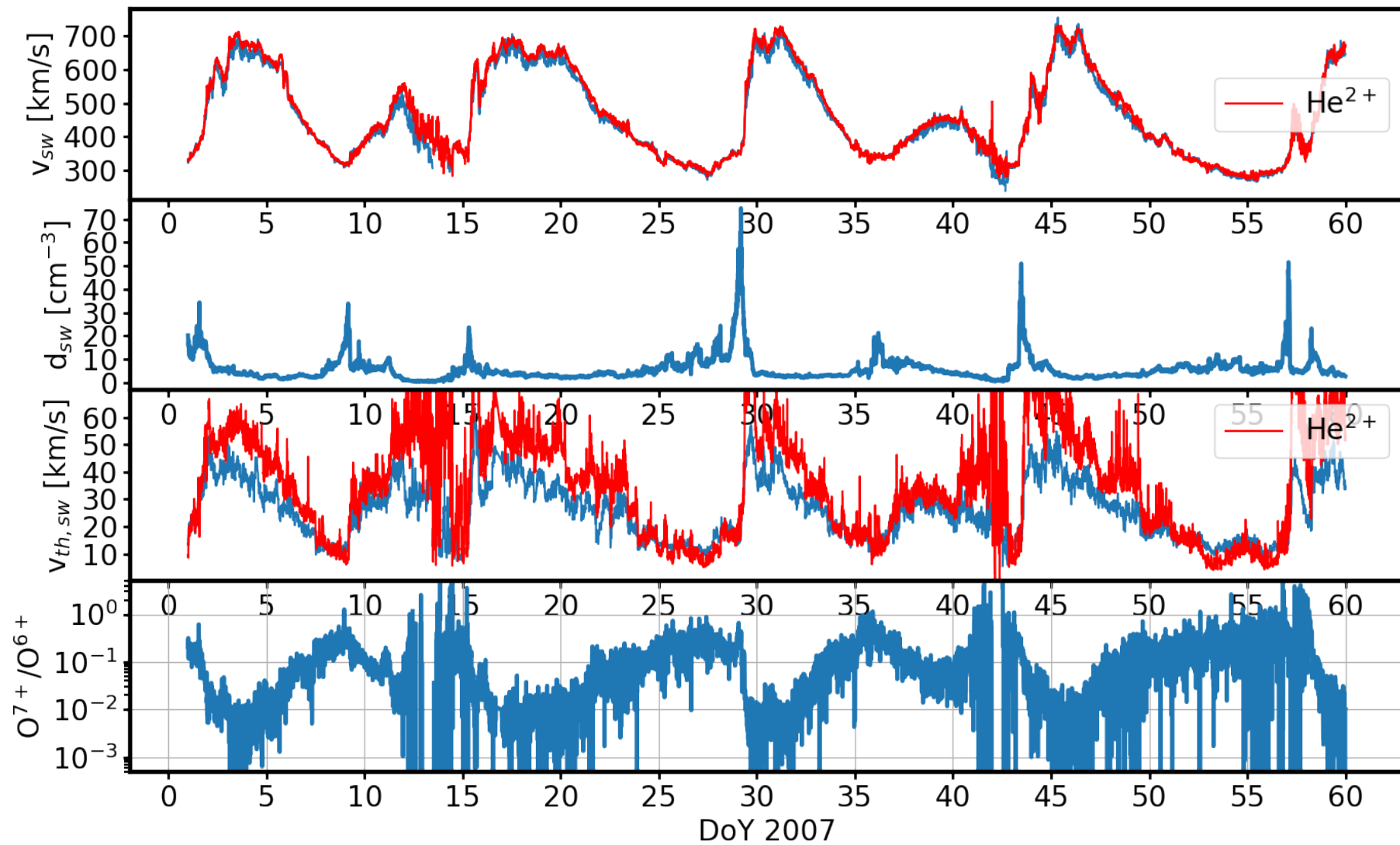
Inner boundary: photospheric magnetic field, taken from synoptic maps (MDI on SOHO)

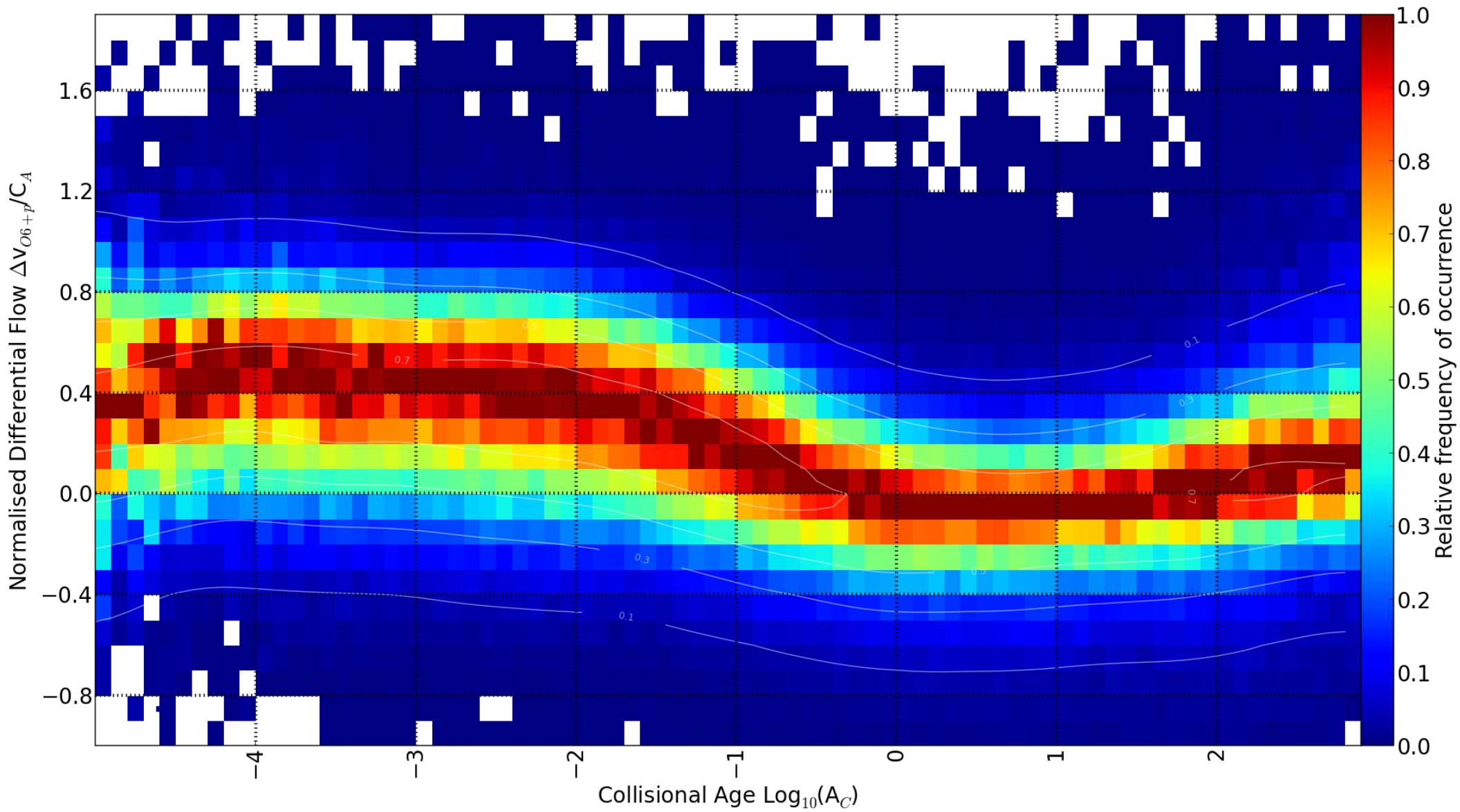
Outer boundary: the source surface at $2.5 R_{\odot}$, here field lines must be strictly radial

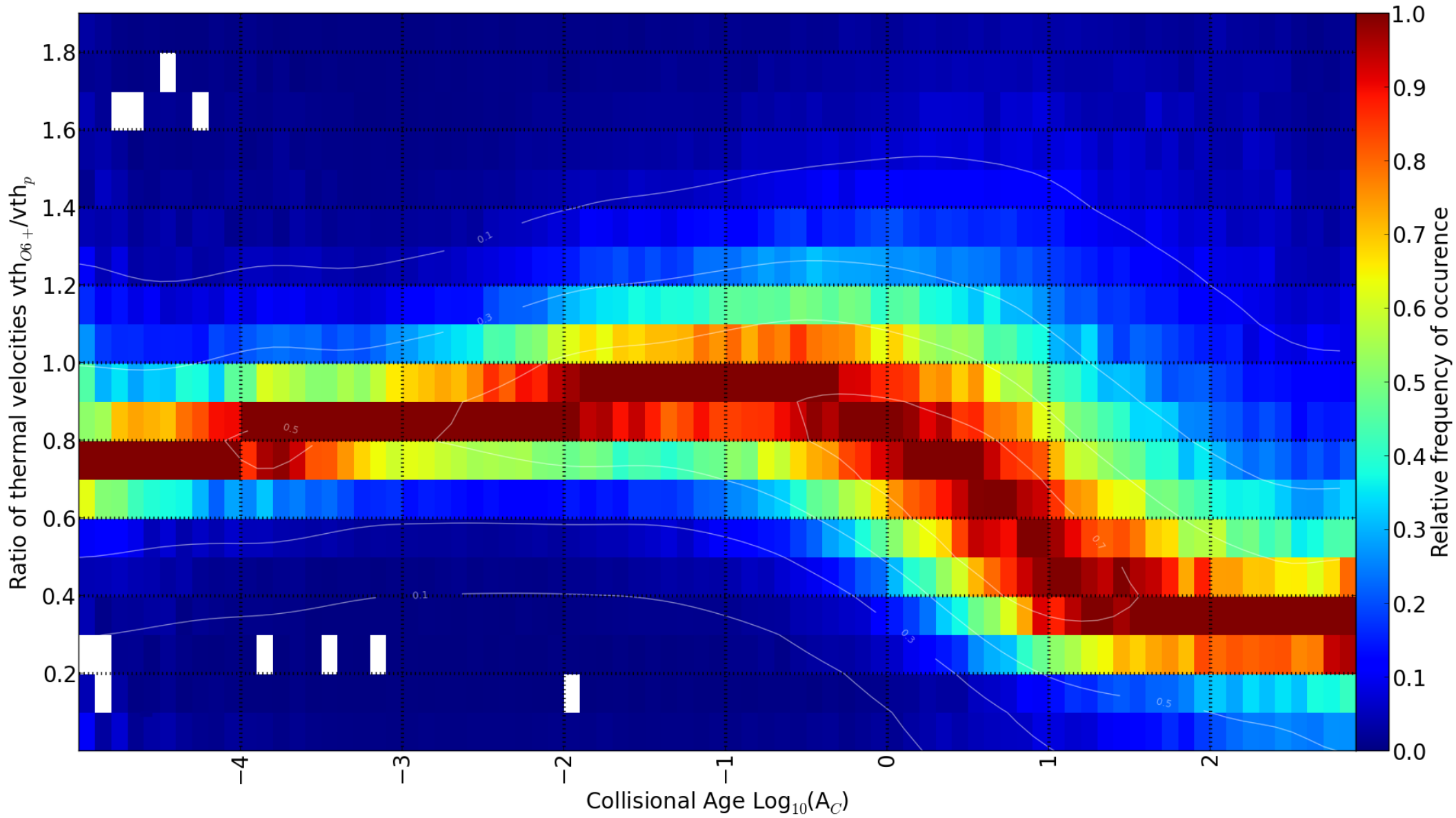




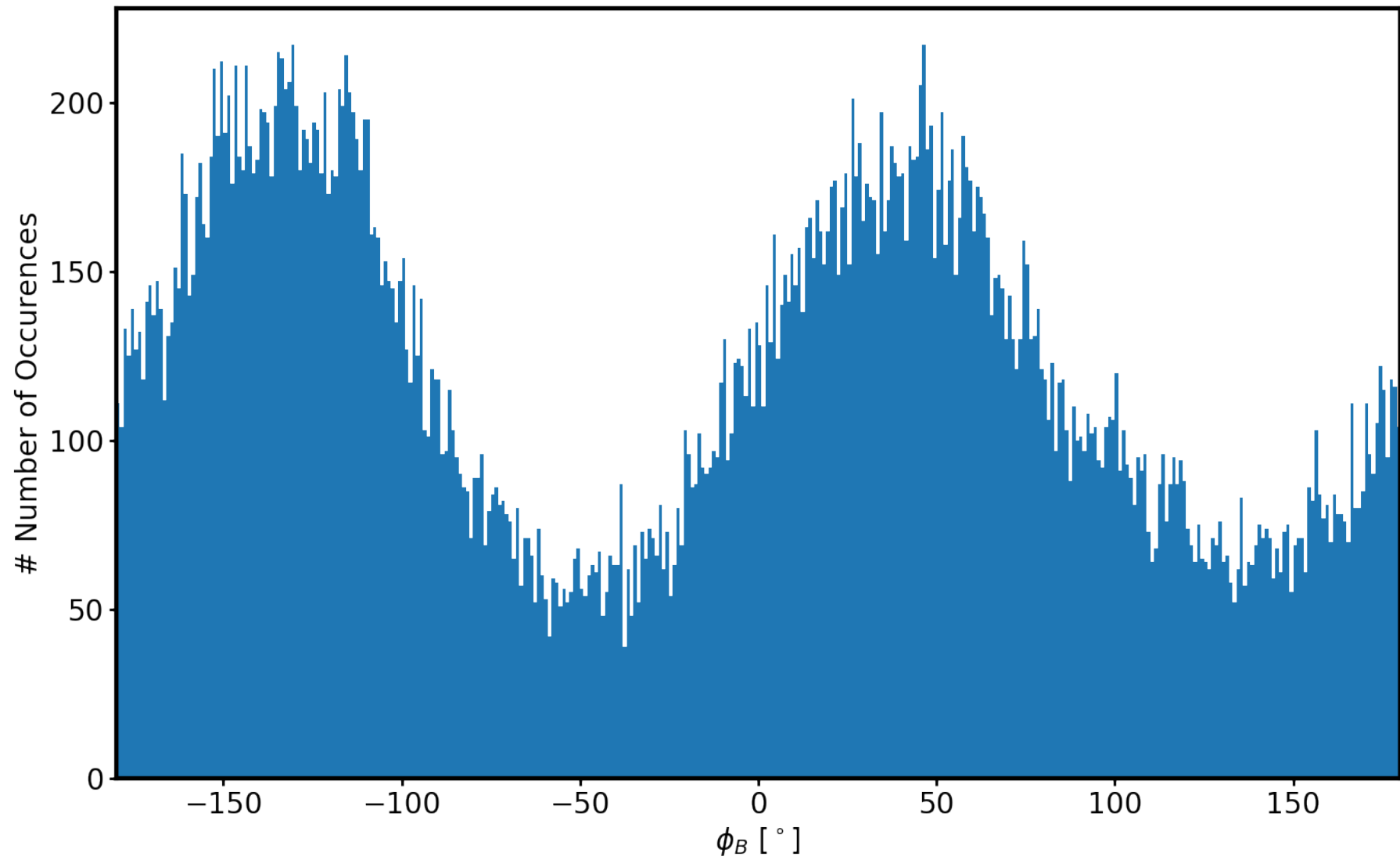




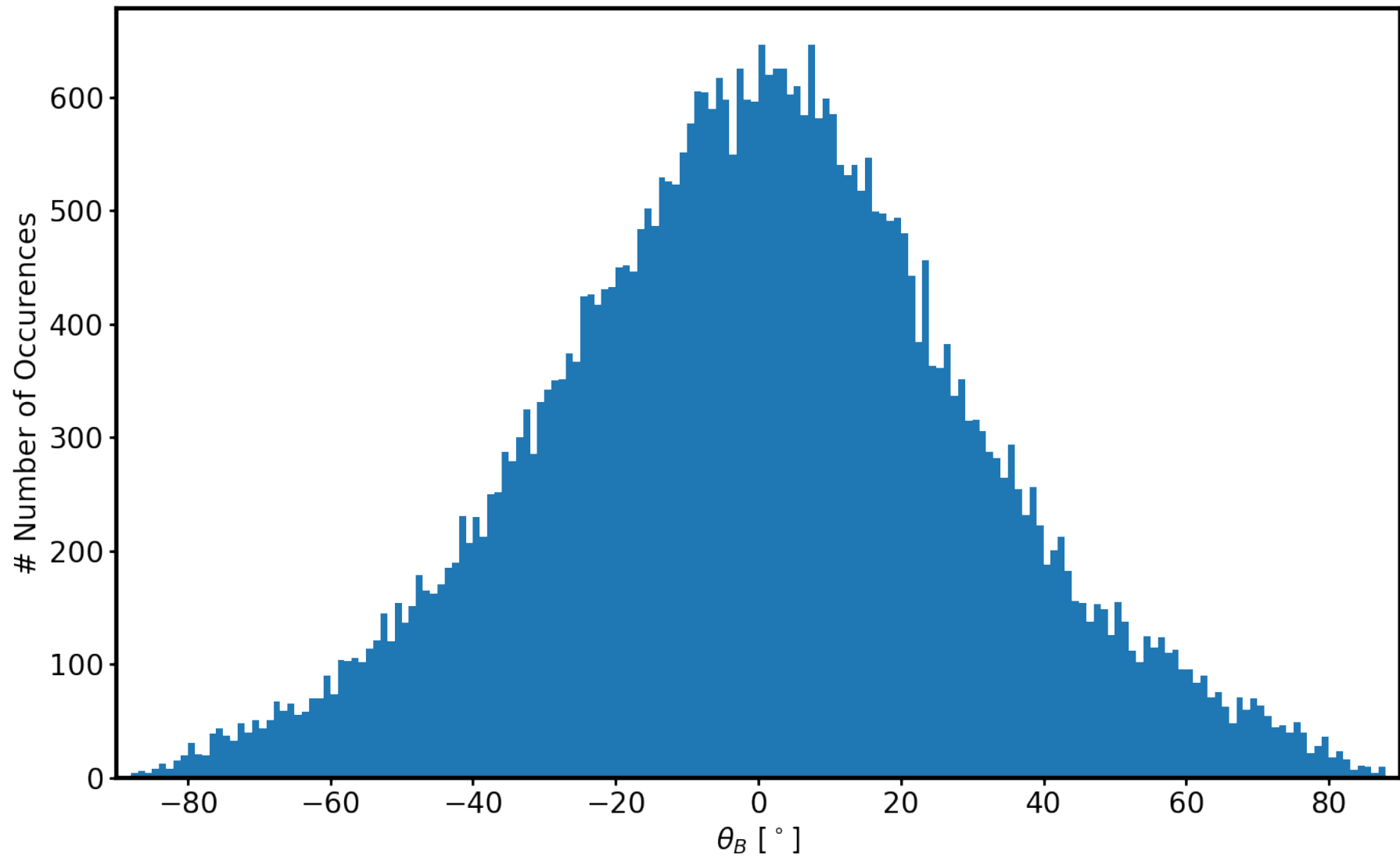


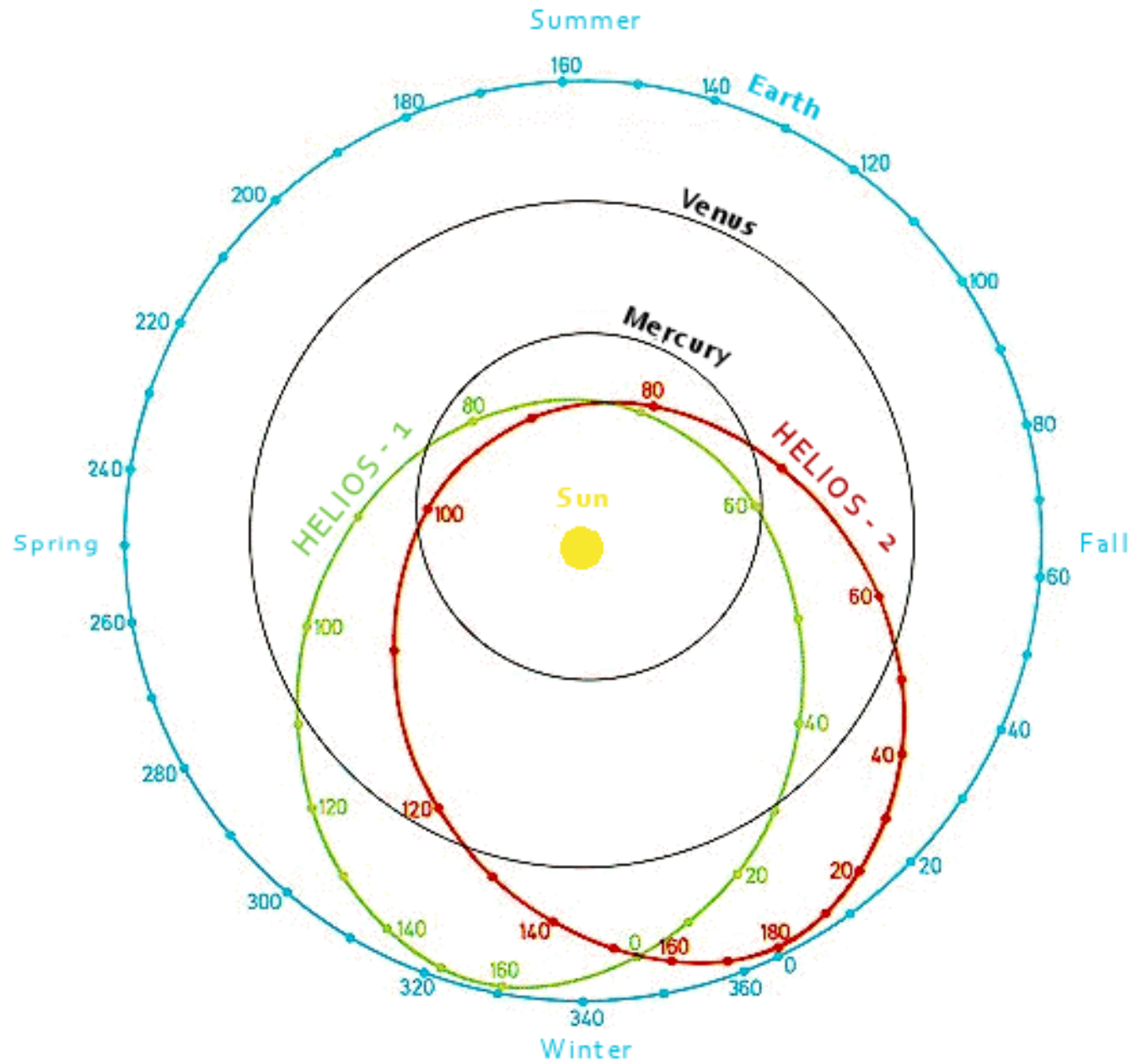


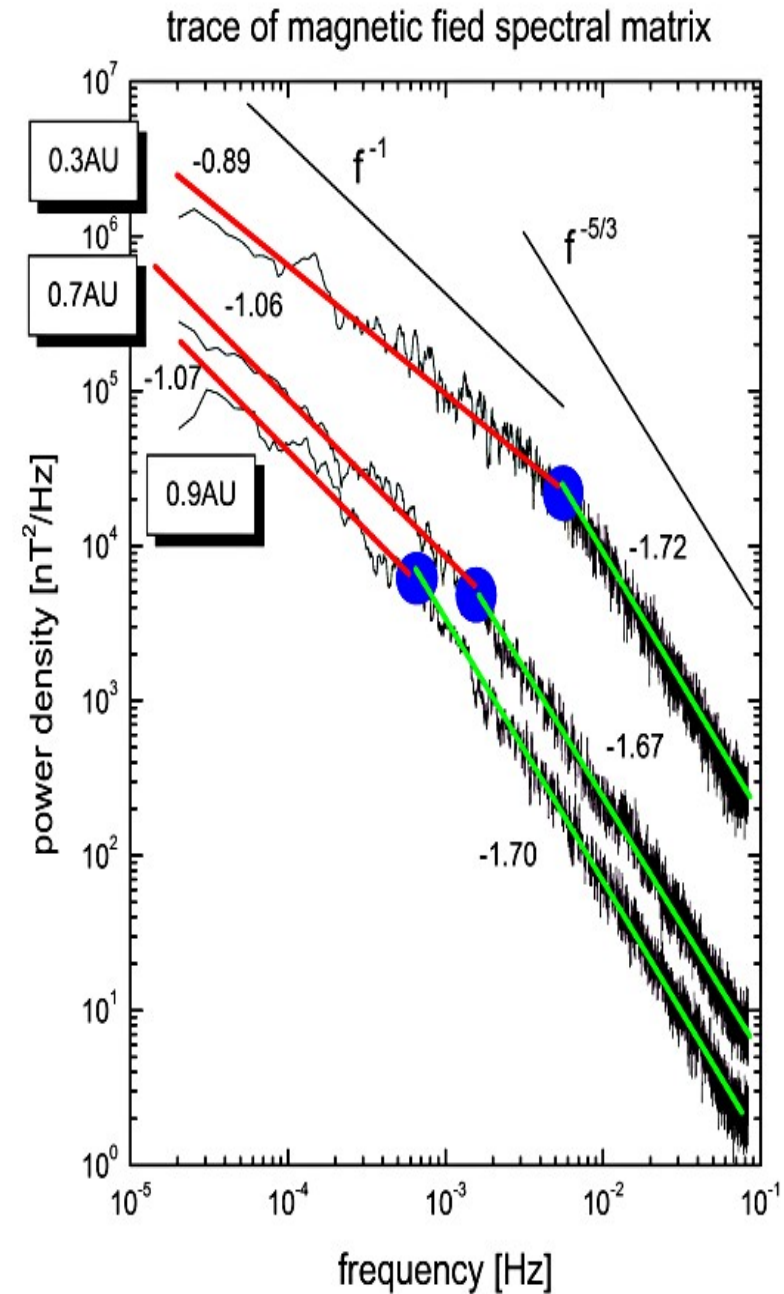
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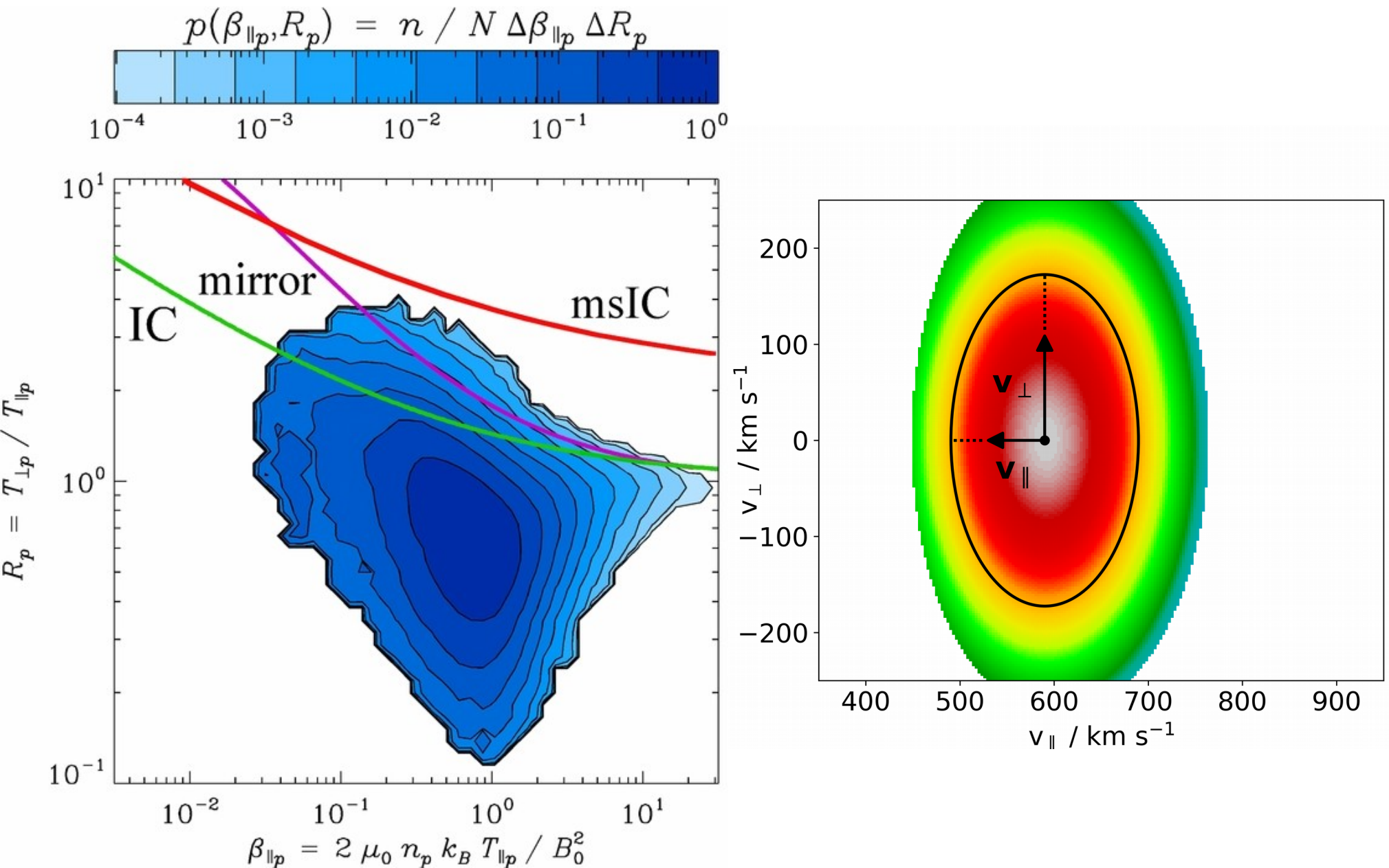


2007

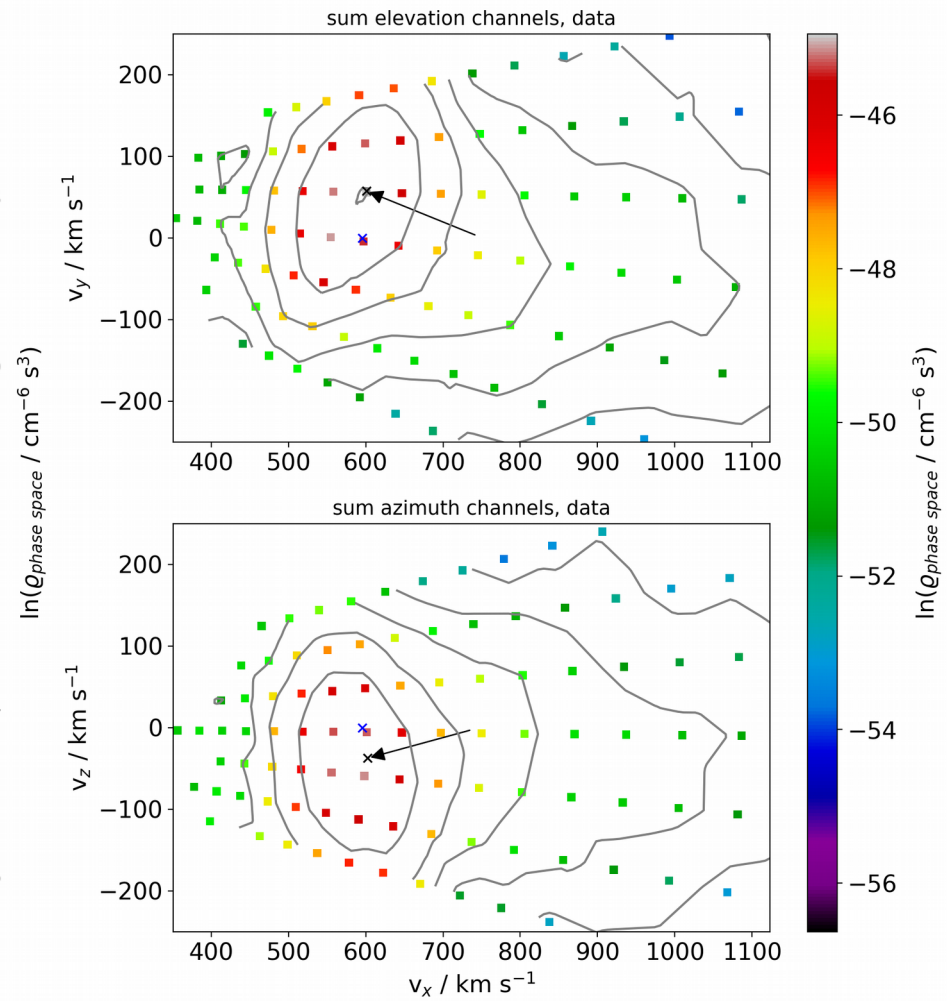
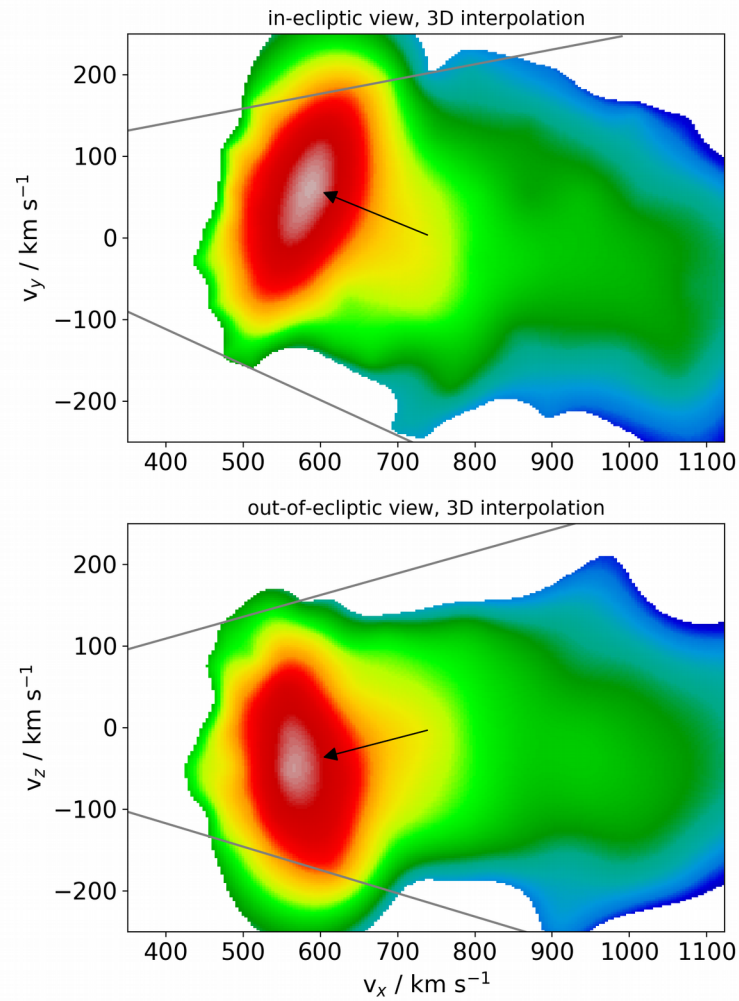


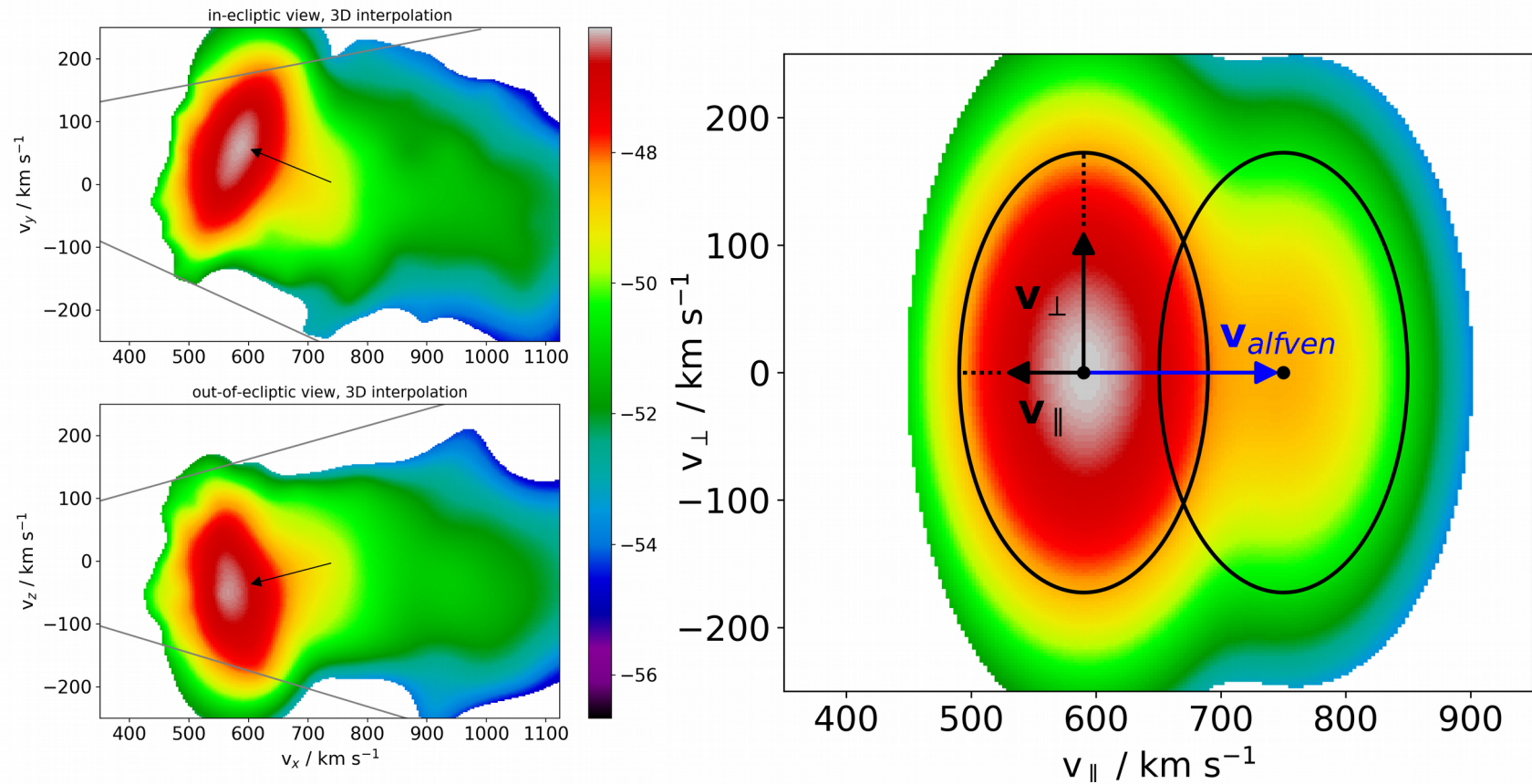






Isenberg et al. 2013





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